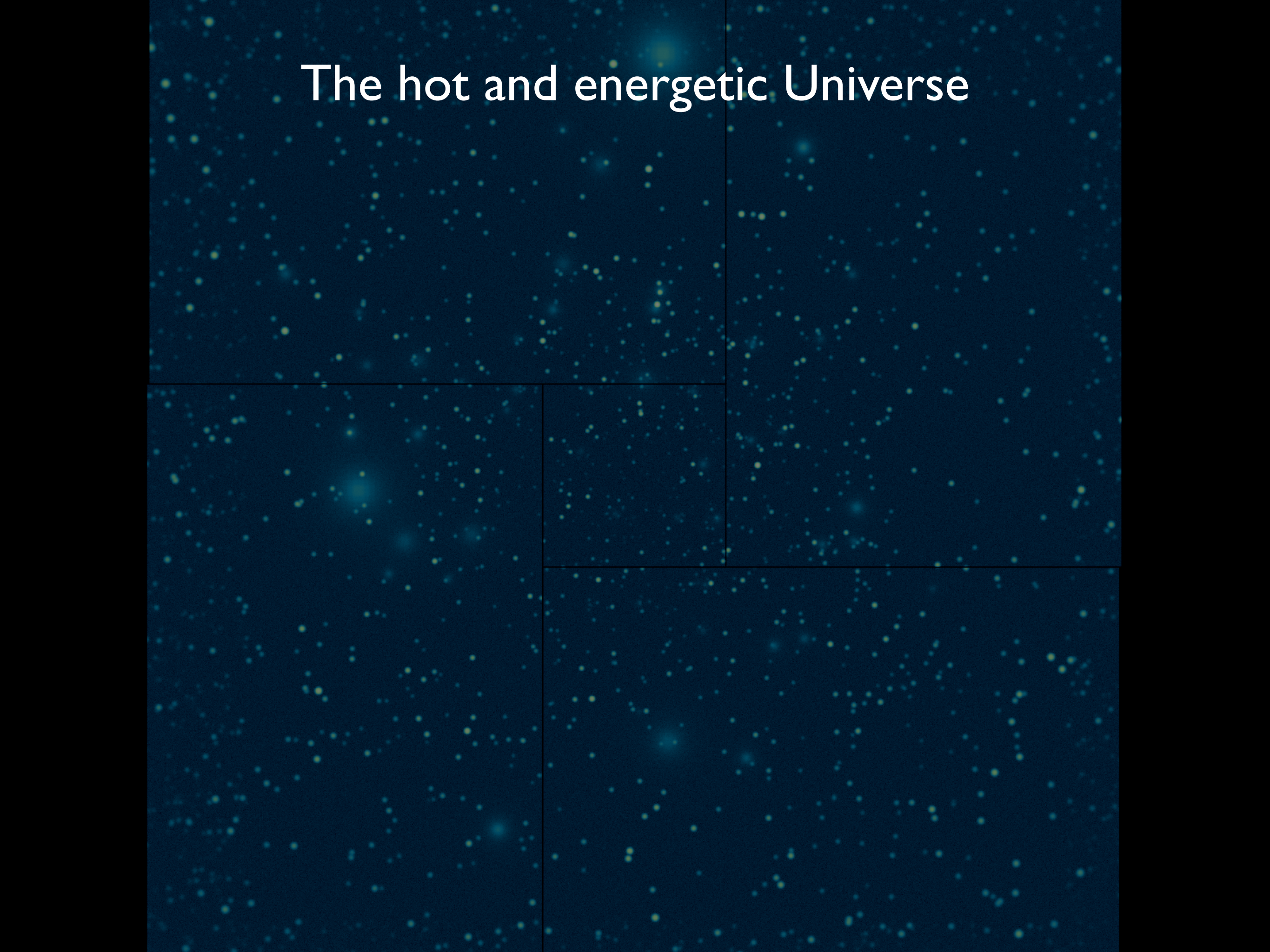


SKA and ATHENA

Gabriel W. Pratt
(SAp, CEA Saclay)

The hot and energetic Universe

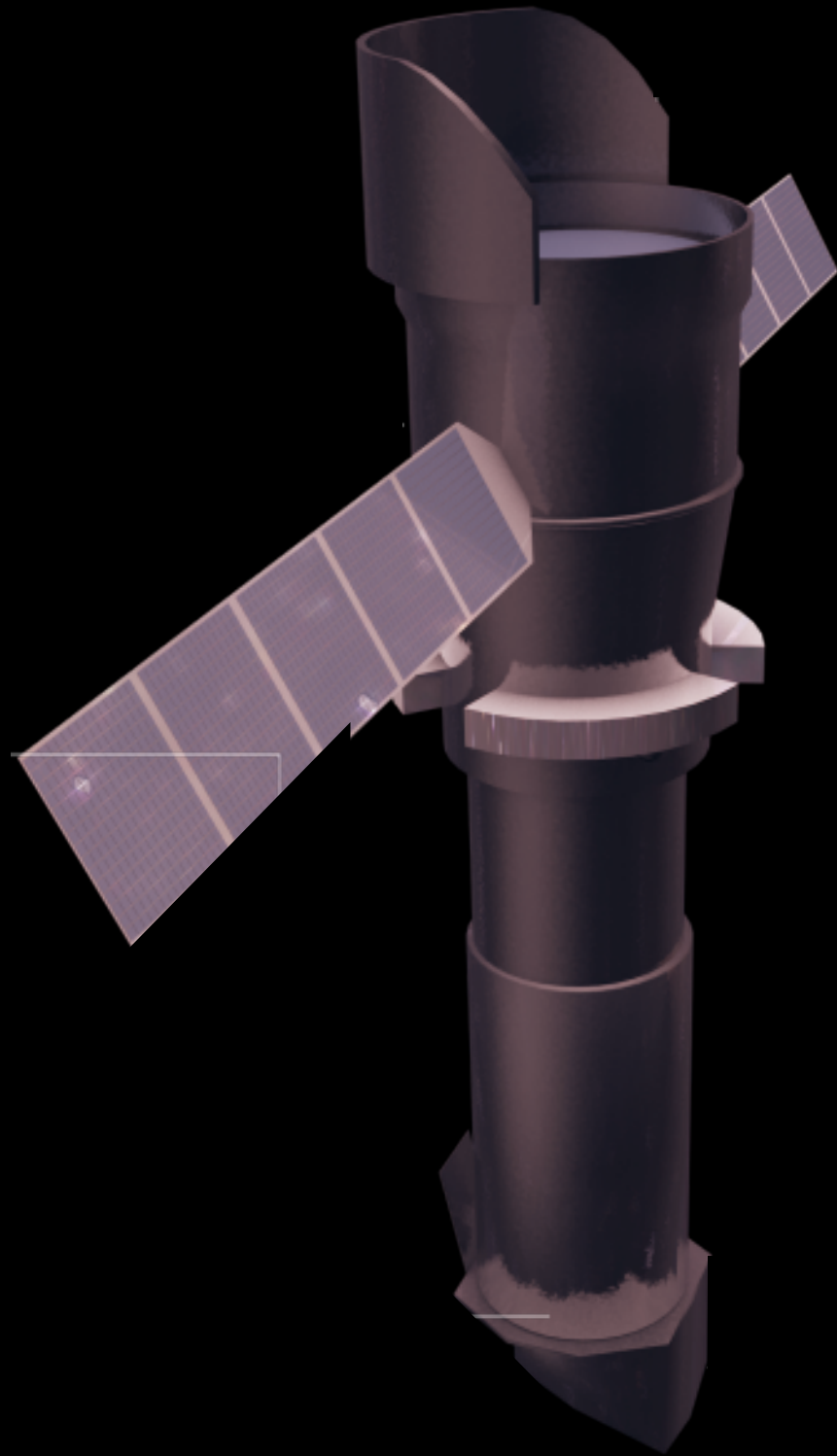
The background of the slide is a deep-field astronomical image, likely from the Hubble Space Telescope, showing a vast field of stars and distant galaxies. The image is divided into four quadrants by a white grid. The top-left quadrant contains the title text. The stars appear as small, bright points of light against the dark cosmic background, with some showing more significant diffraction patterns.

The hot and energetic Universe

How does ordinary matter assemble into the large-scale structures we see today?

How do black holes grow and shape the Universe?

The Advanced Telescope for High ENergy Astrophysics



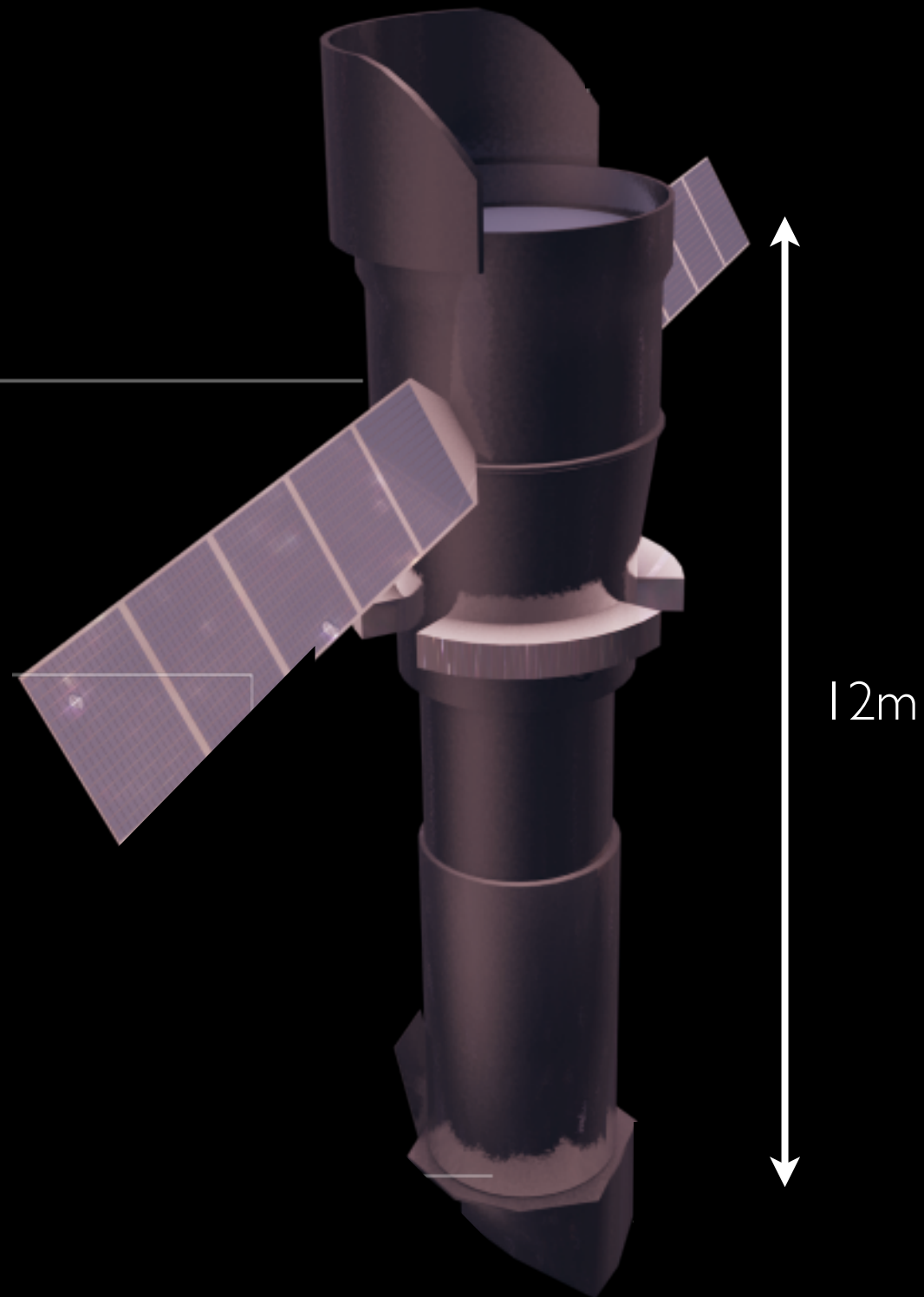
The Advanced Telescope for High ENergy Astrophysics

L2 orbit Ariane V

Mass < 5100 kg

Power 2500 W

5 year mission



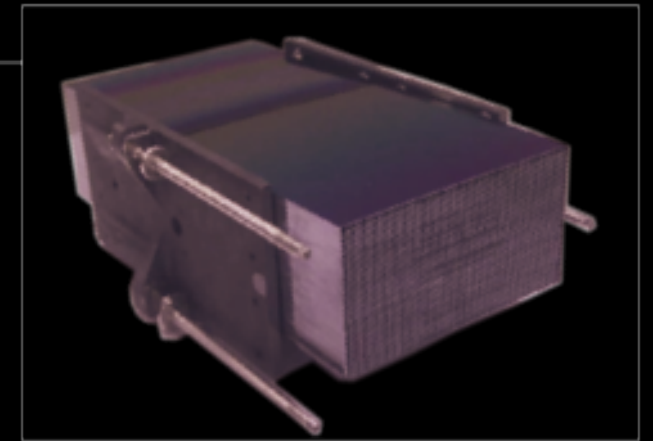
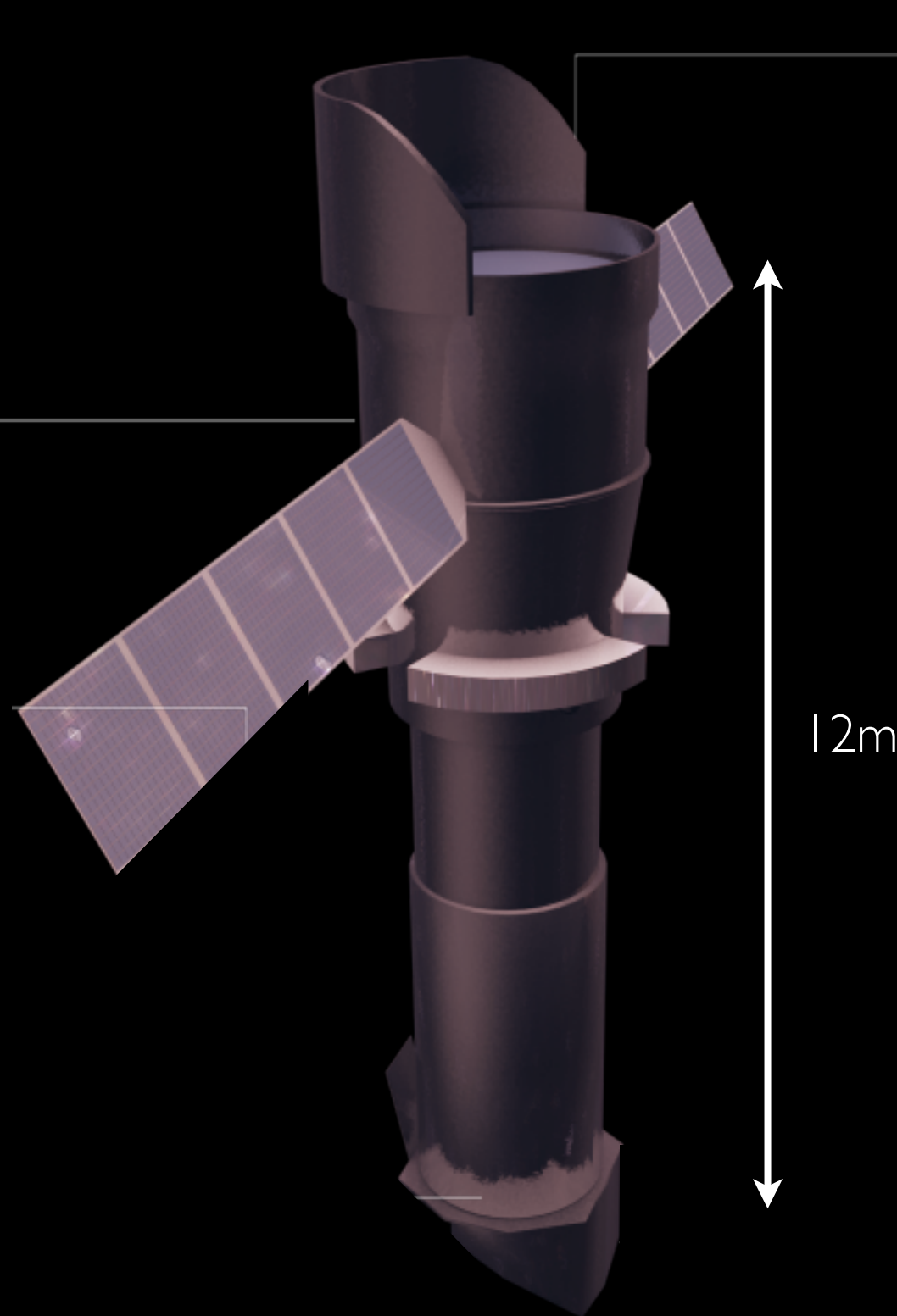
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Silicon Pore Optics:

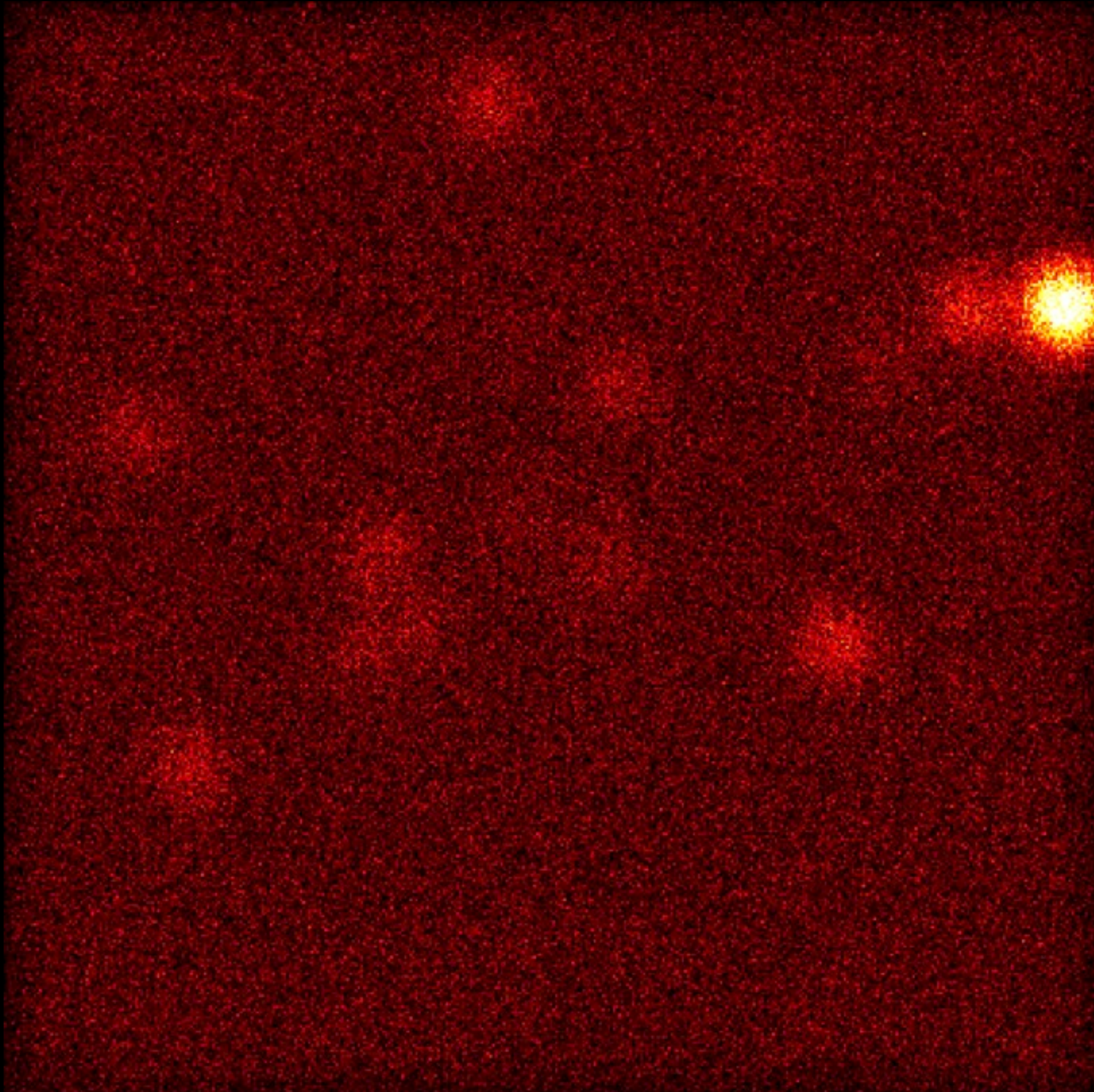
2 m² at 1 keV

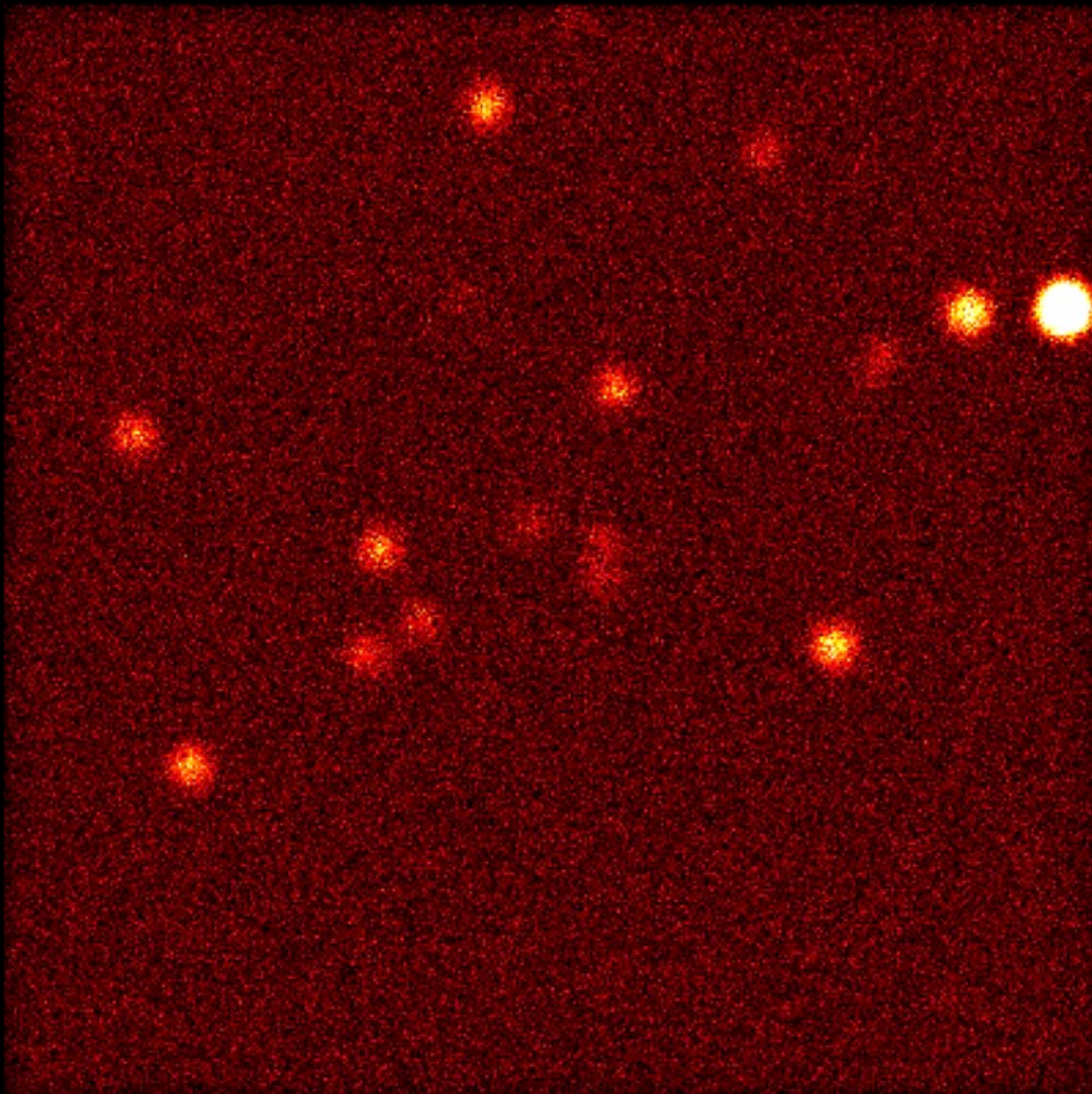
5 arcsec HEW

Focal length: 12 m

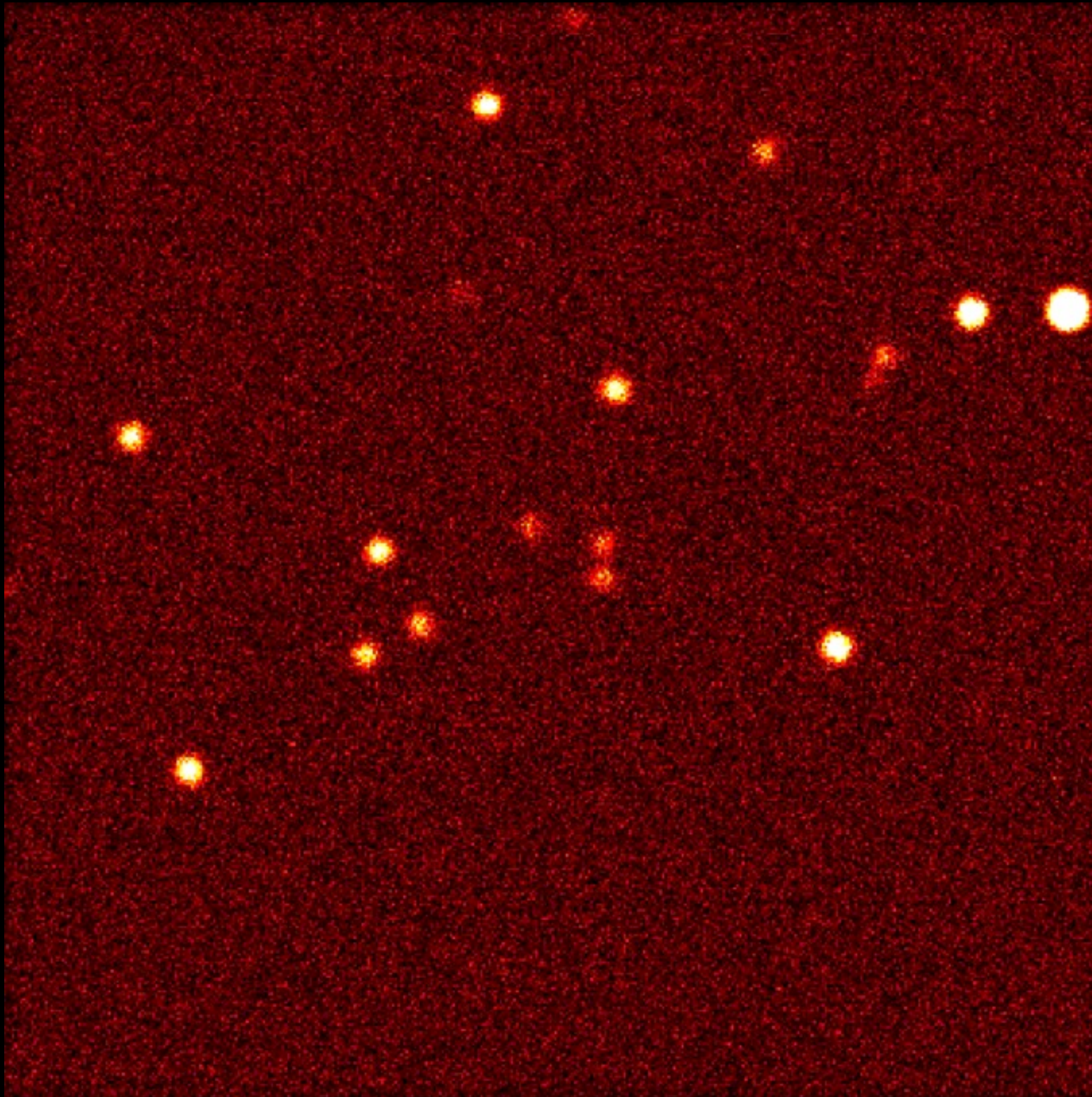
Sensitivity: 3×10^{-17} erg cm⁻² s⁻¹

10''





5''
(requirement)



3''
(goal)

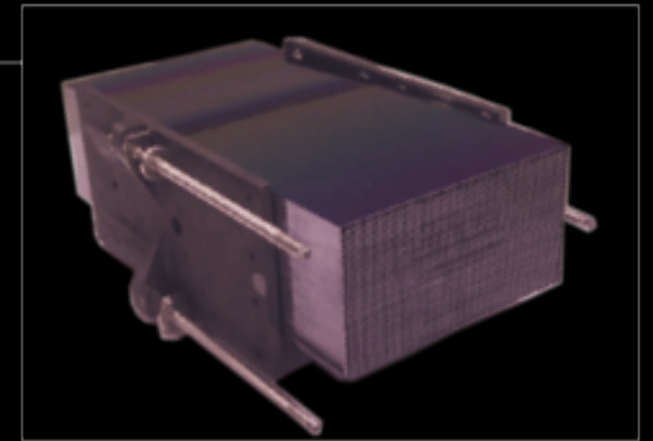
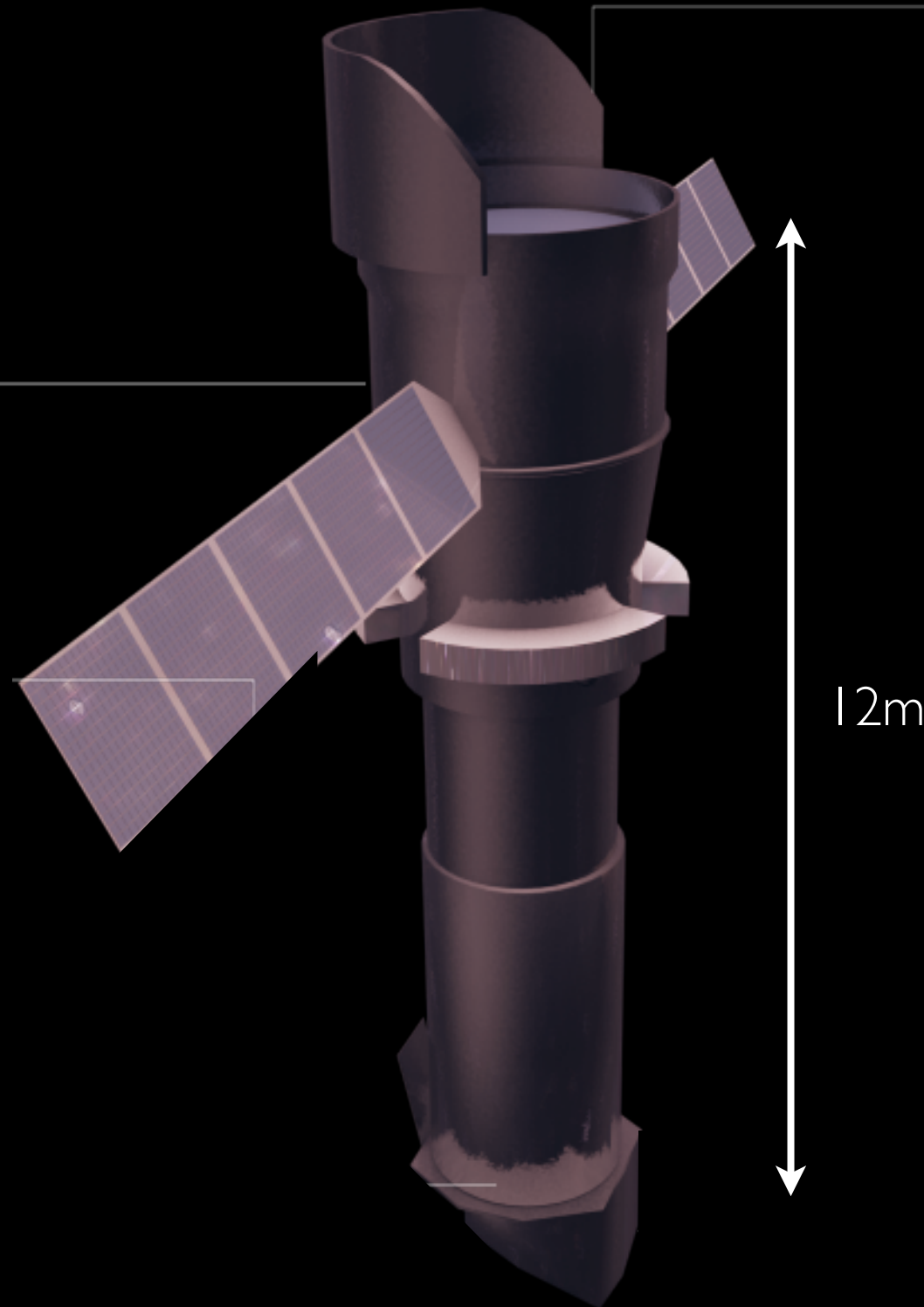
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2 m² at 1 keV

5 arcsec HEW

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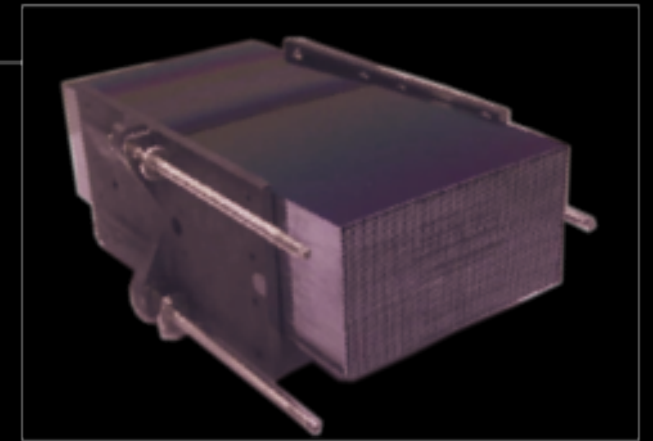
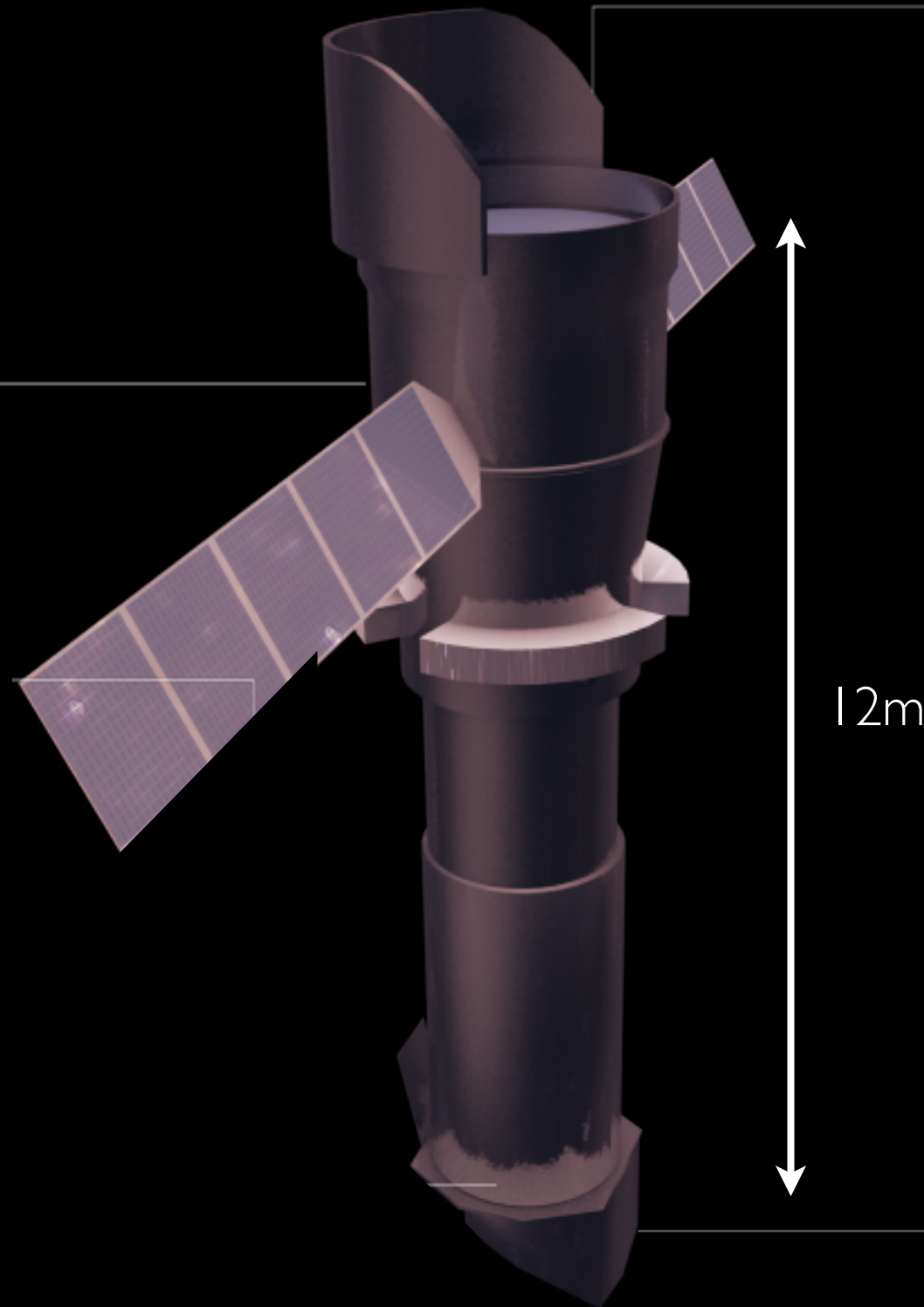
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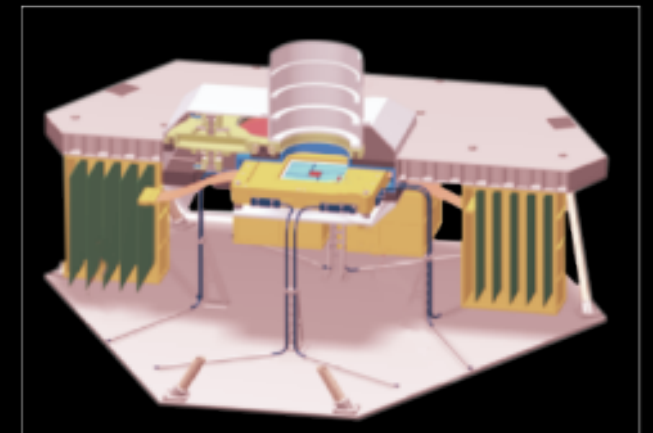
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5 arcsec HEW

Focal length: 12 m

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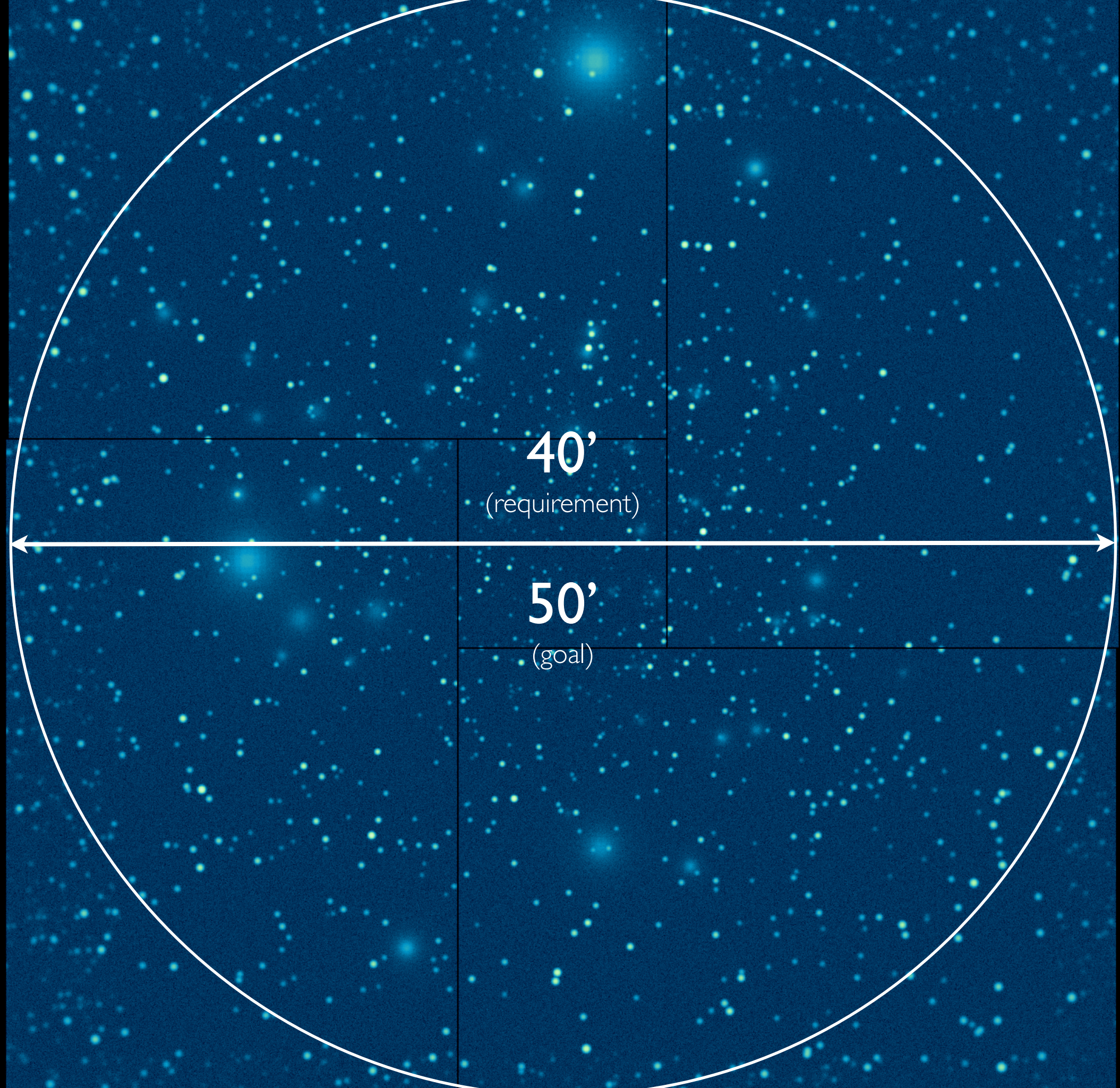


Wide Field Imager:

ΔE: 125 eV

Field of View: 40 arcmin

High countrate capability



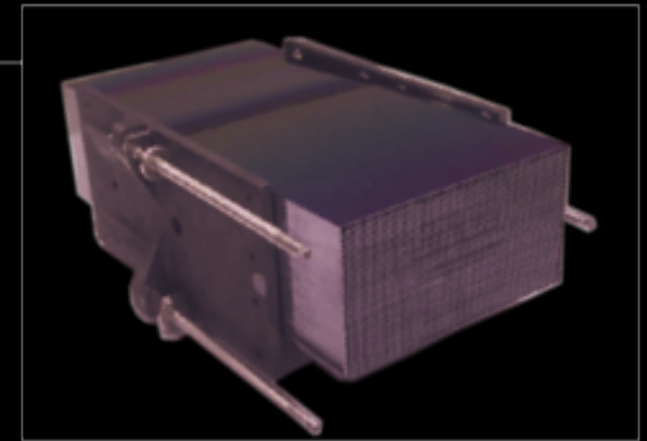
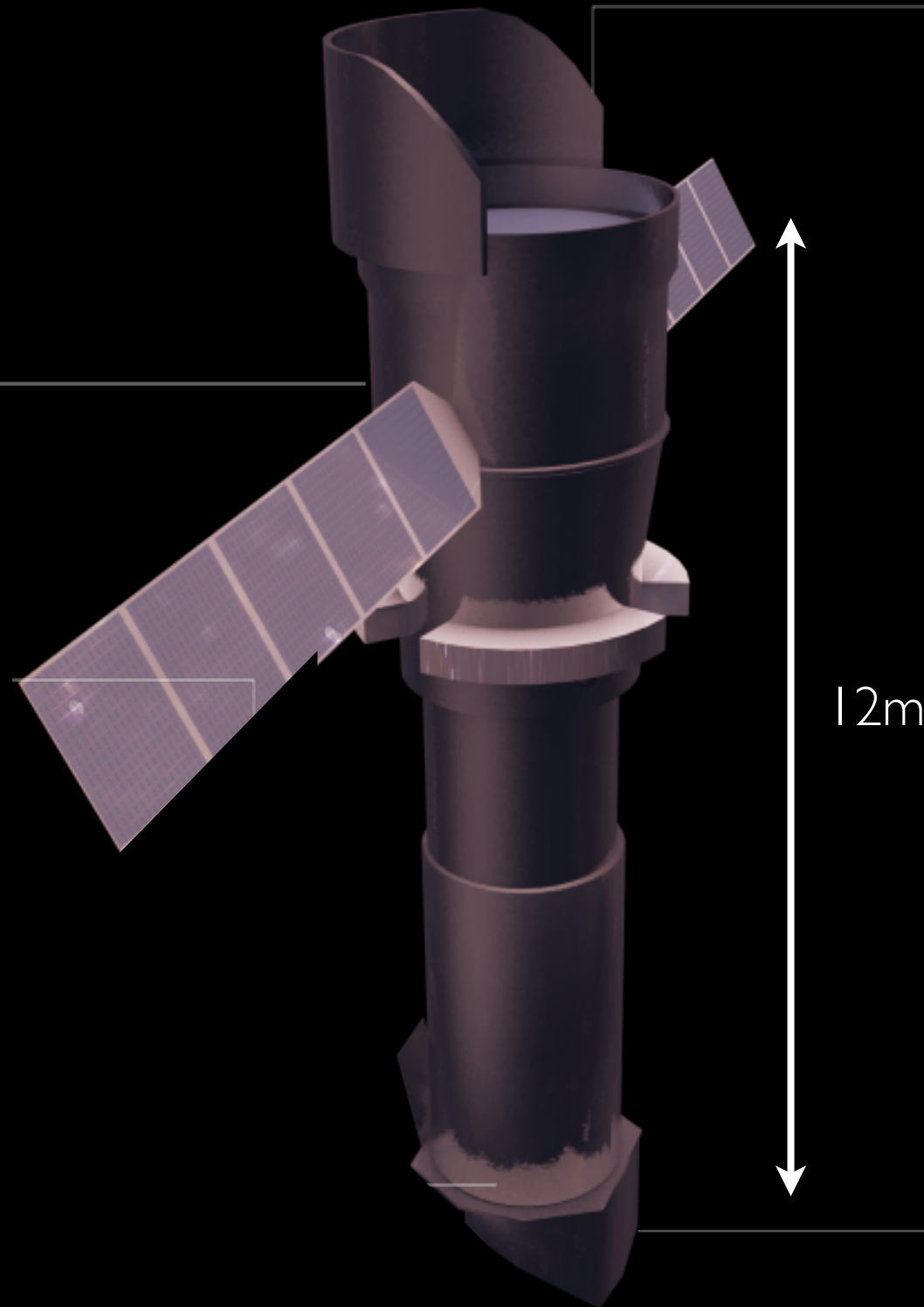
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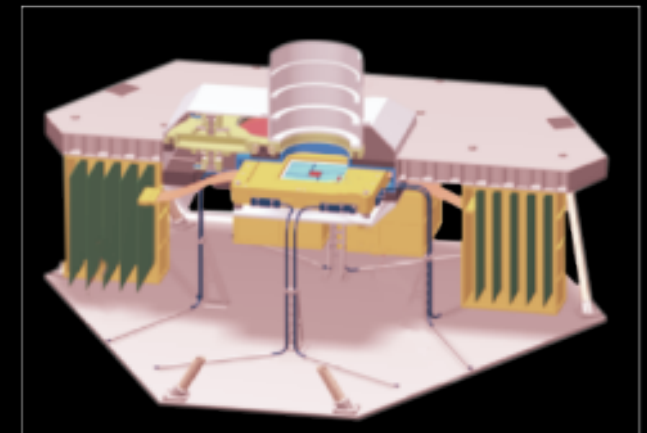
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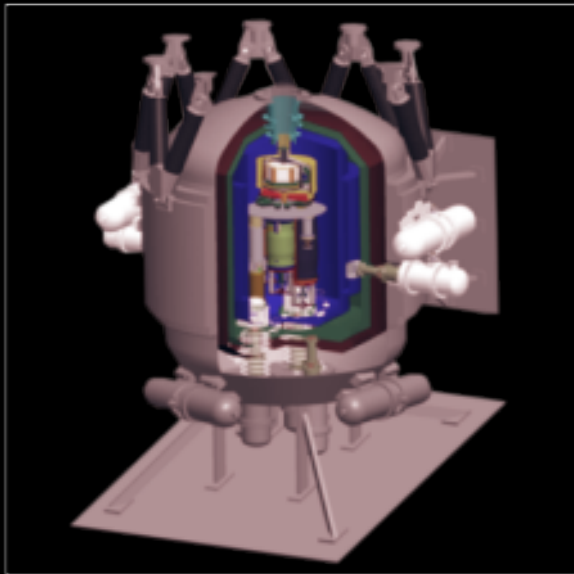
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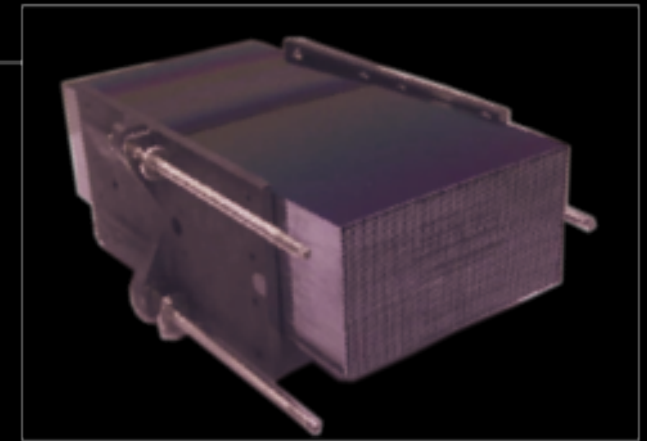
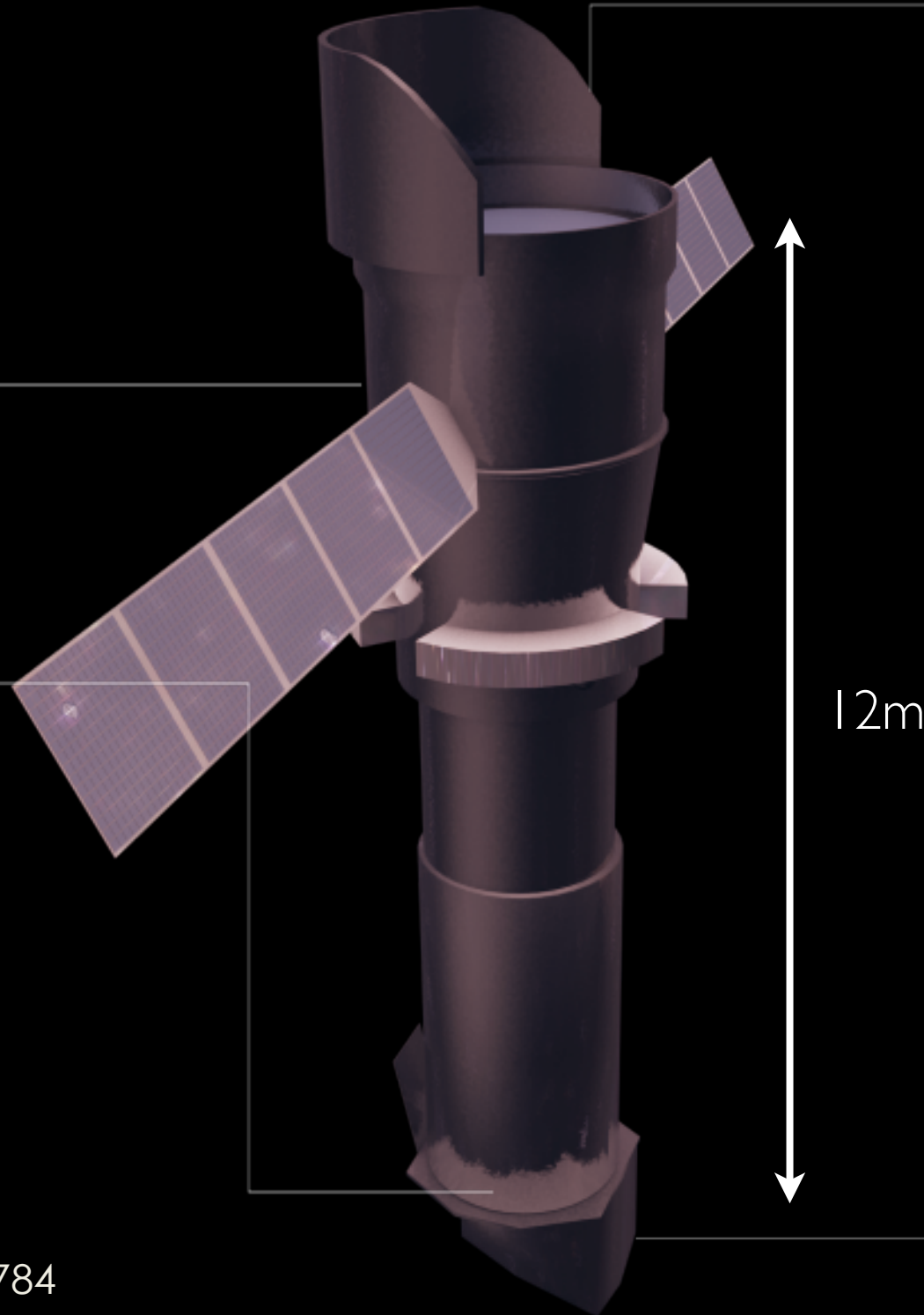


X-ray Integral Field Unit:

ΔE : 2.5 eV

Field of View: 5 arcmin

Operating temp: 50 mk



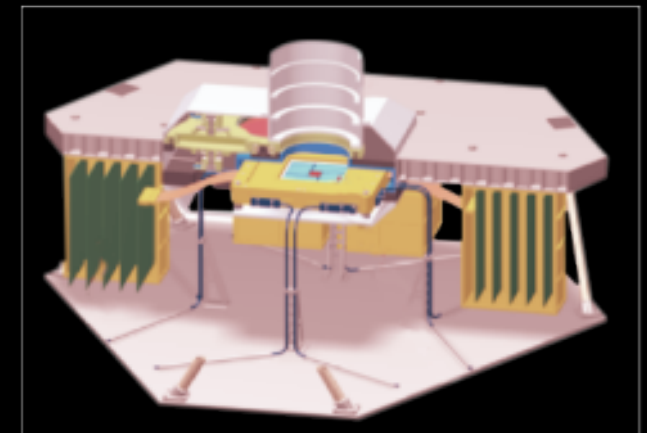
Silicon Pore Optics:

2 m² at 1 keV

5 arcsec HEW

Focal length: 12 m

Sensitivity: $3 \cdot 10^{-17}$ erg cm⁻² s⁻¹



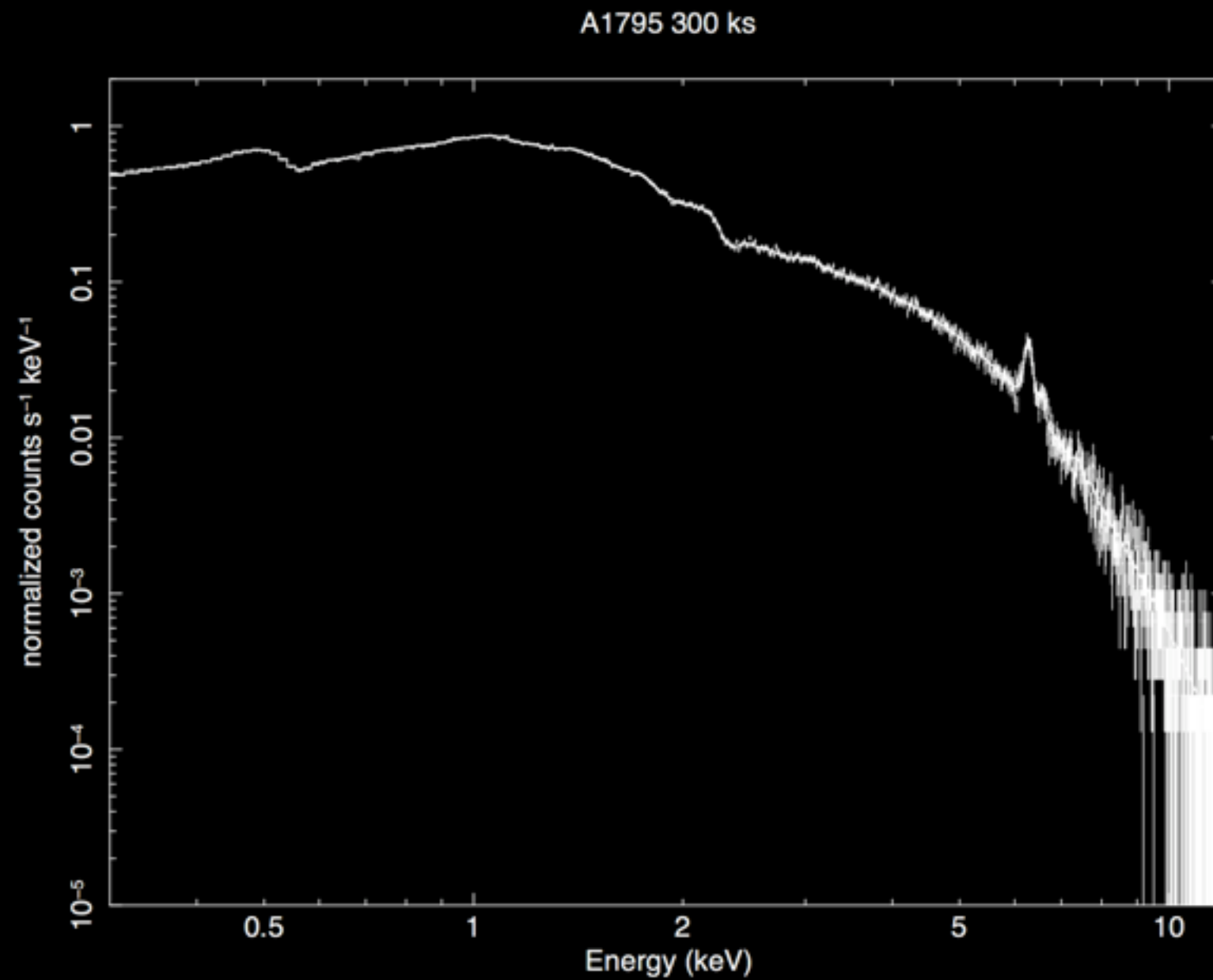
Wide Field Imager:

ΔE : 125 eV

Field of View: 40 arcmin

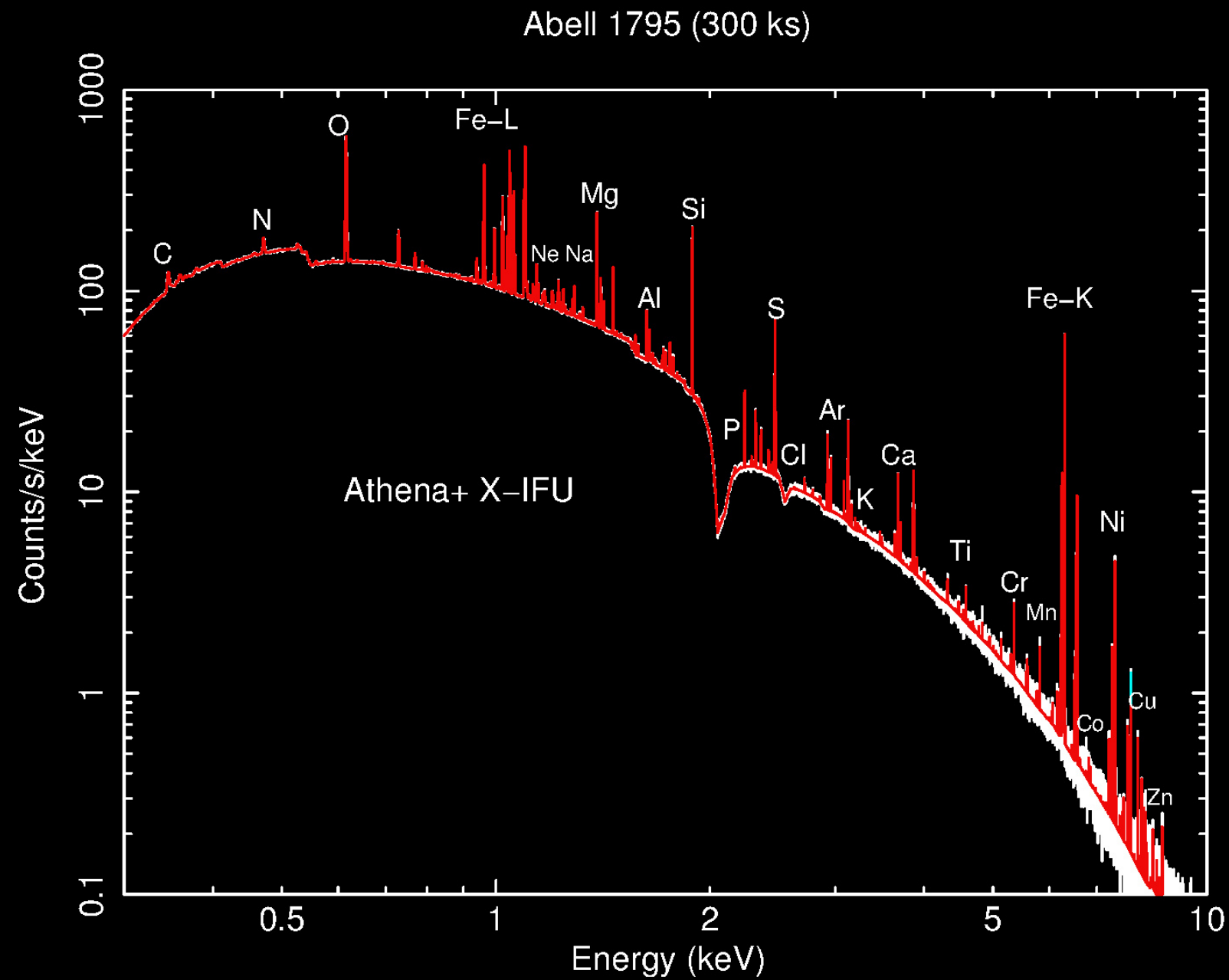
High countrate capability

XIFU



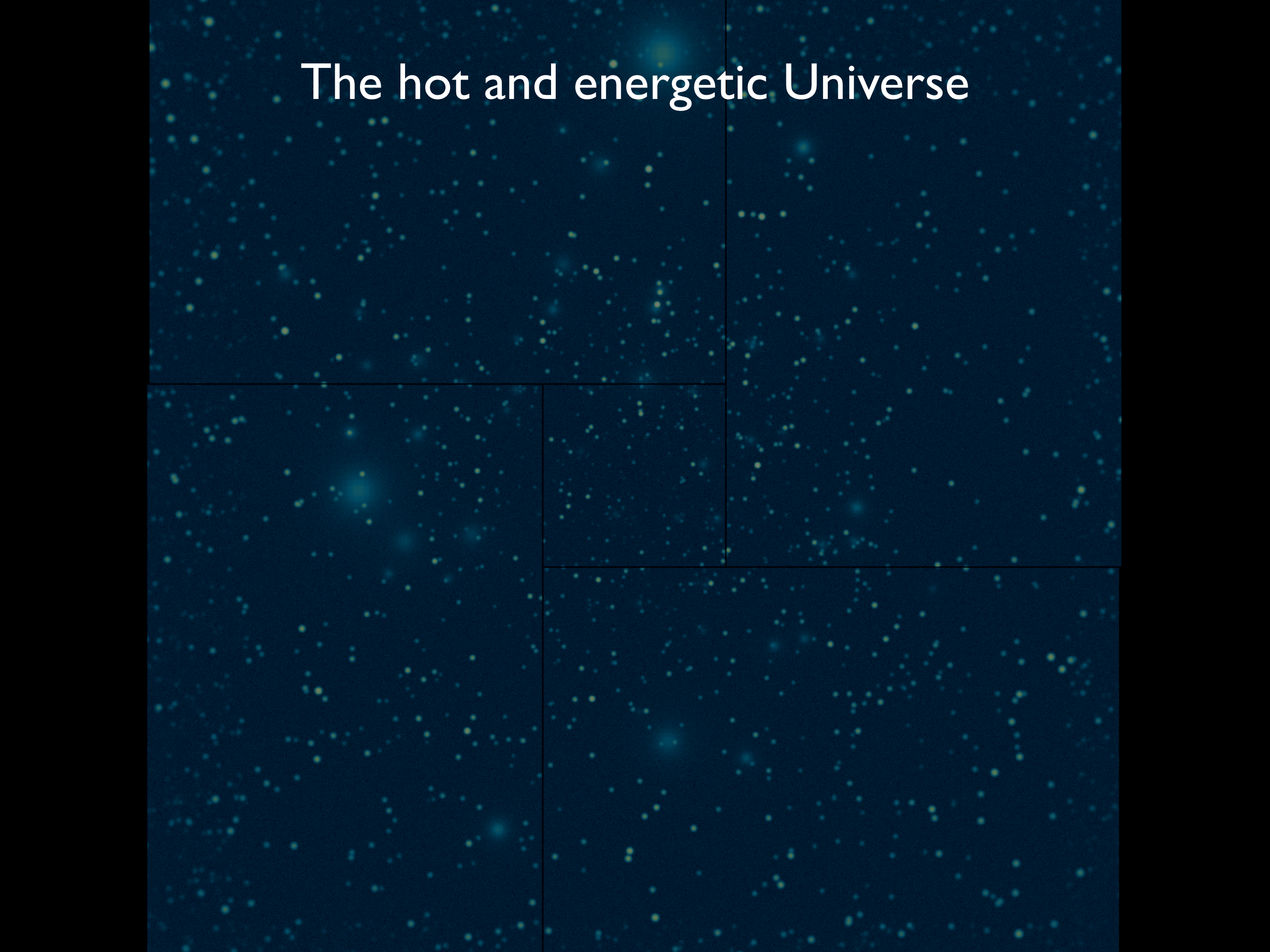
But smaller field of view

XIFU



But smaller field of view

The hot and energetic Universe

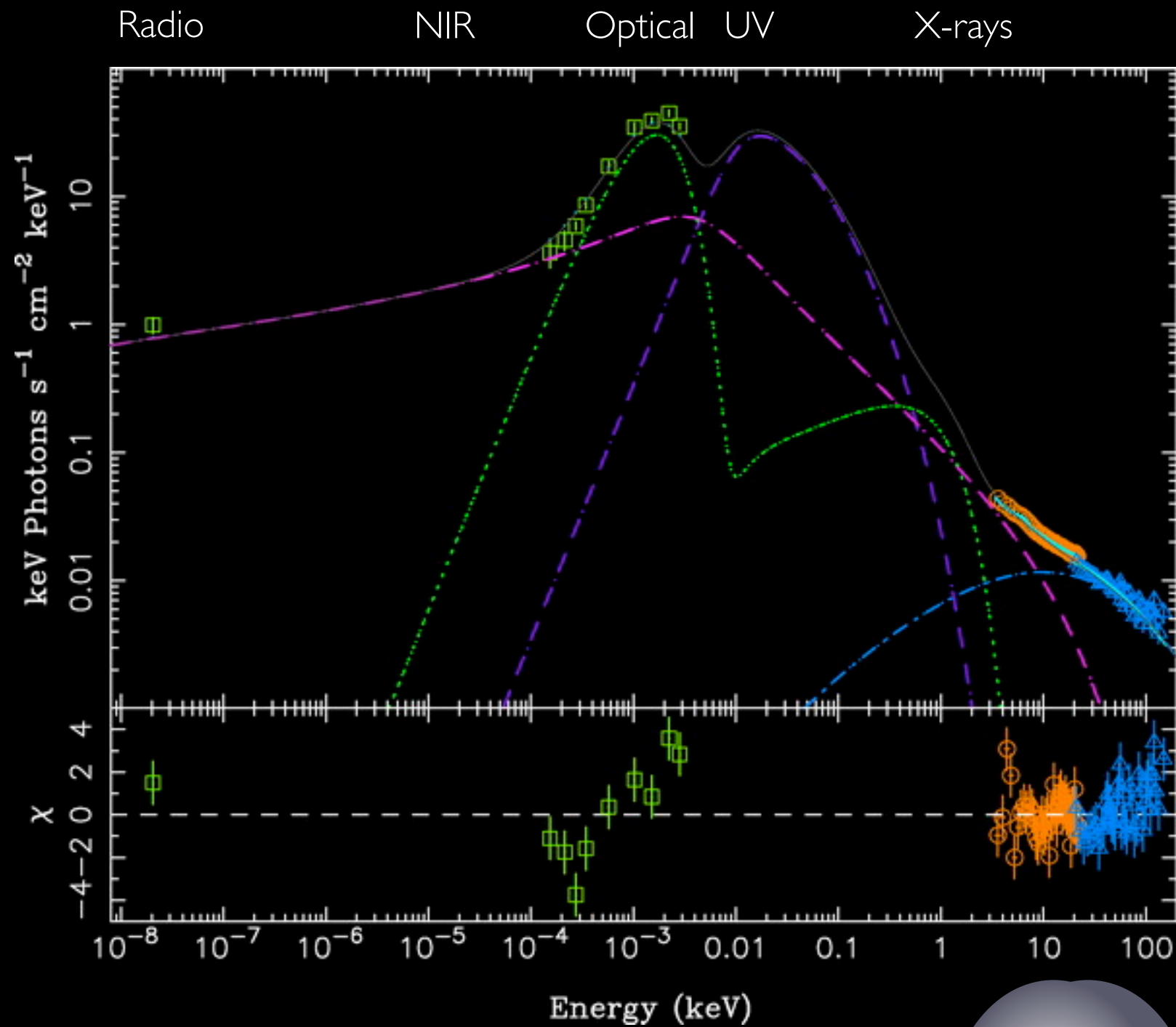
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The hot and energetic Universe

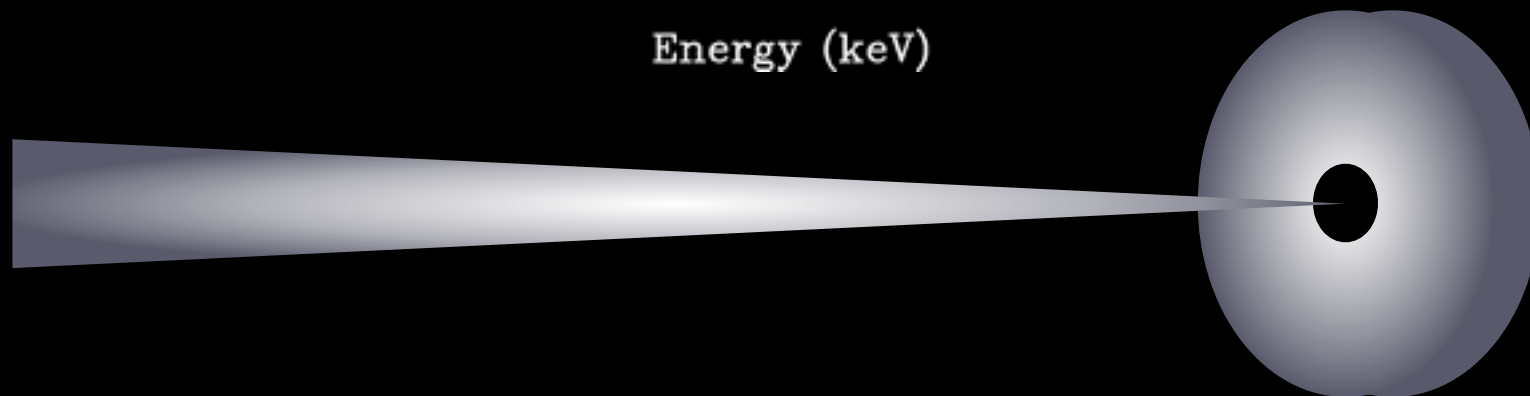
- ▶ Collecting area
- ▶ Spatial resolution
- ▶ Field of View (WFI)
- ▶ Spectral resolution (XIFU)
- ▶ (timing resolution [μ s WFI])

Synergies with SKA and pathfinders (a few examples)

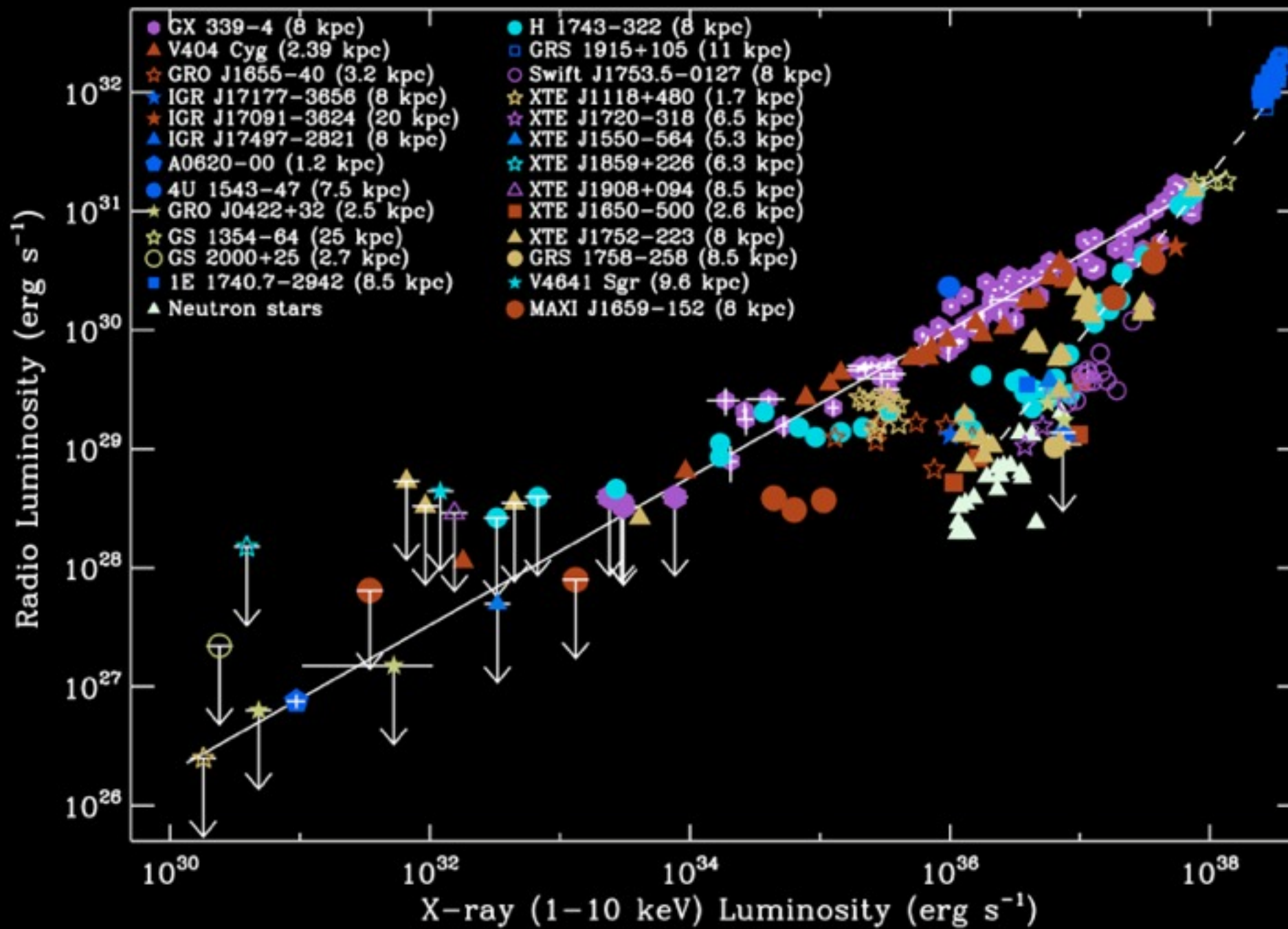
Microquasars



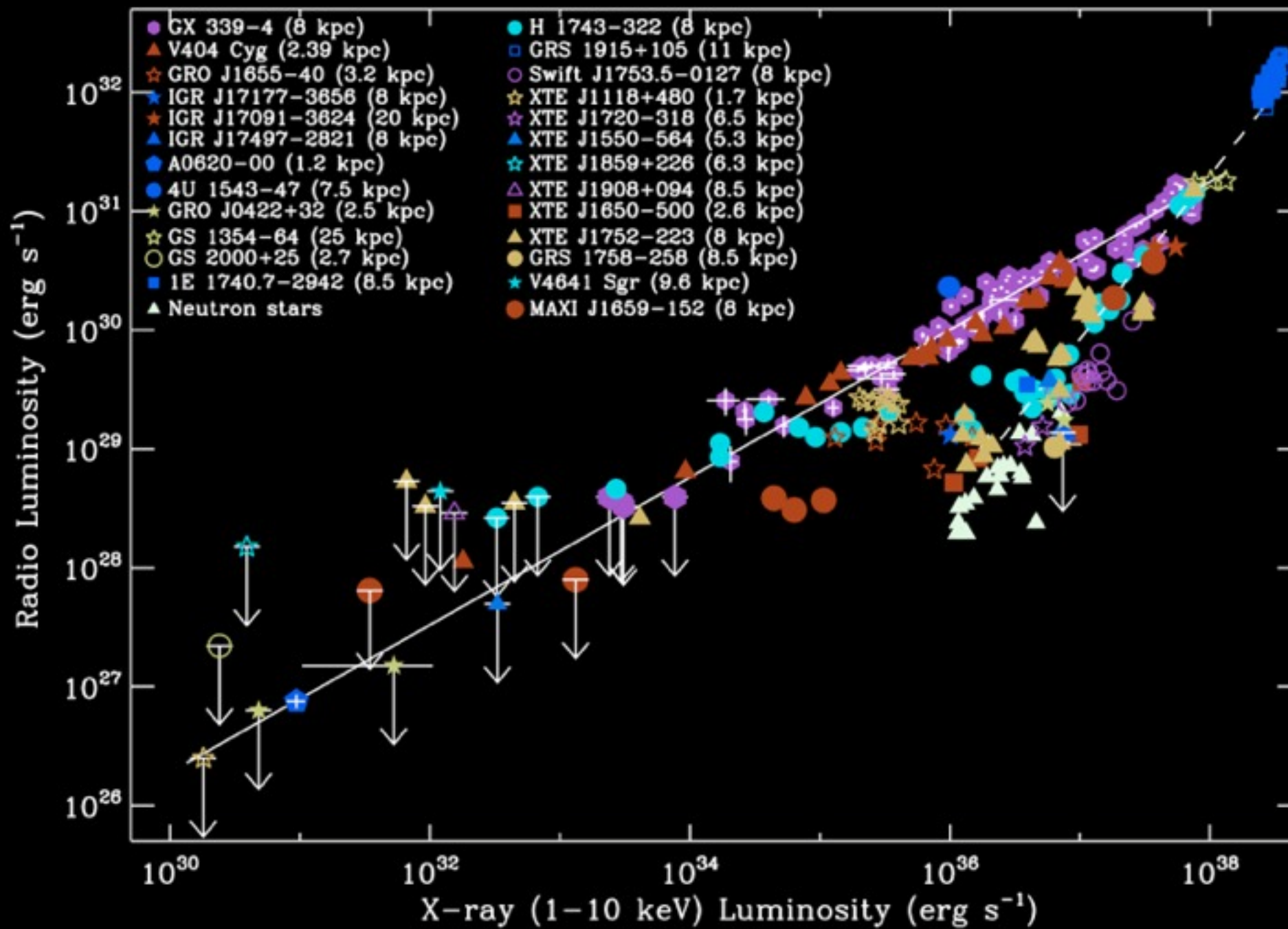
preshock synchrotron
 postshock synchrotron
 SSC + disk external
 Compton
 disk + companion star
 blackbody



Microquasars

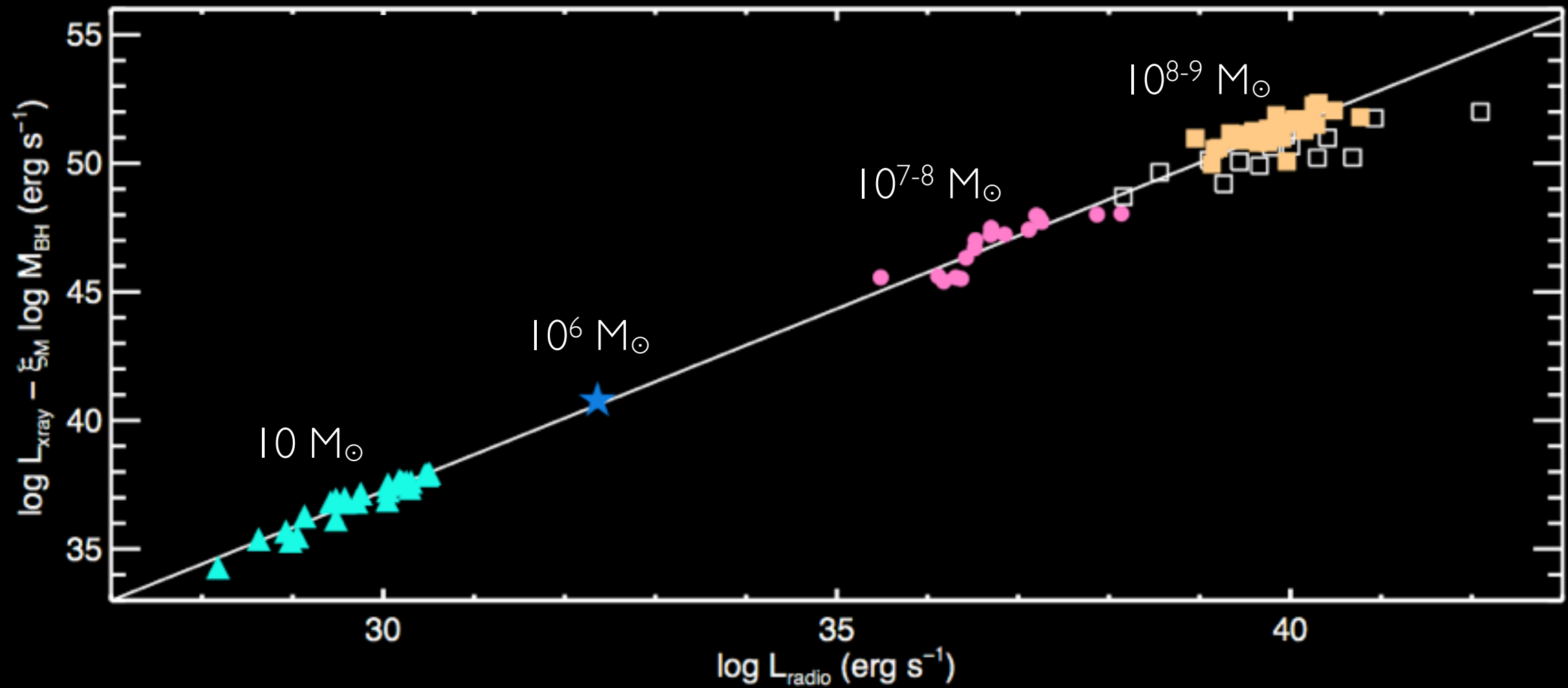


Microquasars



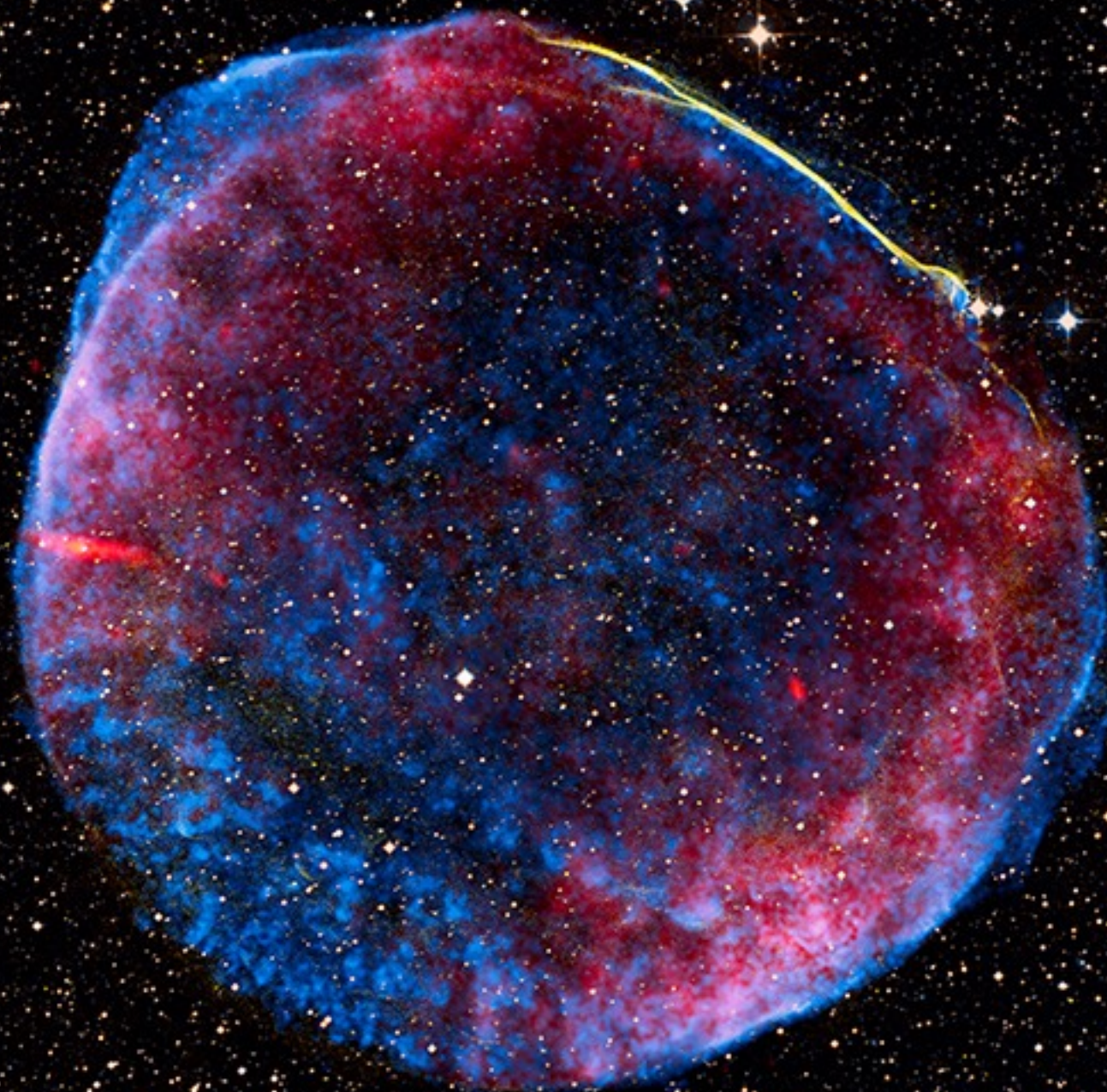
Corbel et al 2013

Black hole L_X - L_R fundamental plane



Plotkin et al 2012

SN 1006

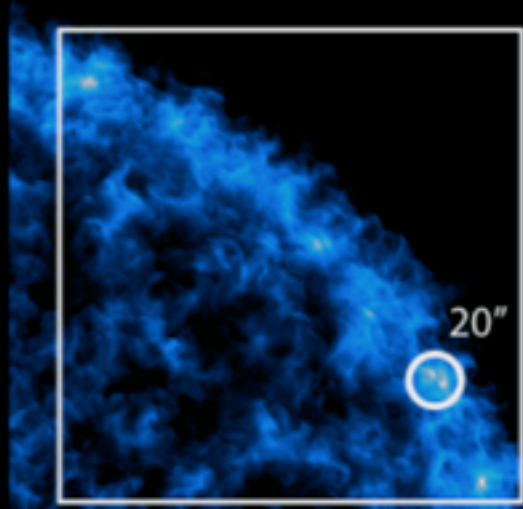


X-ray
Radio
optical

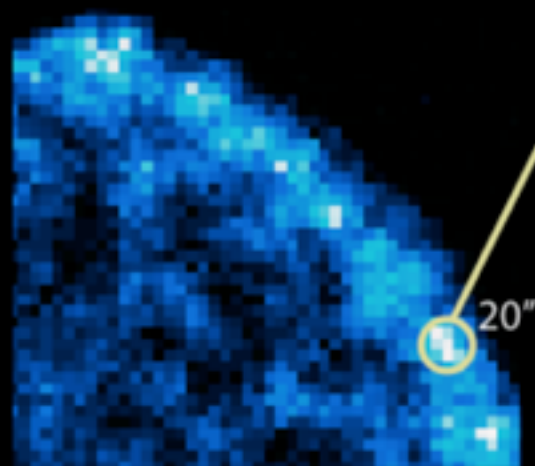
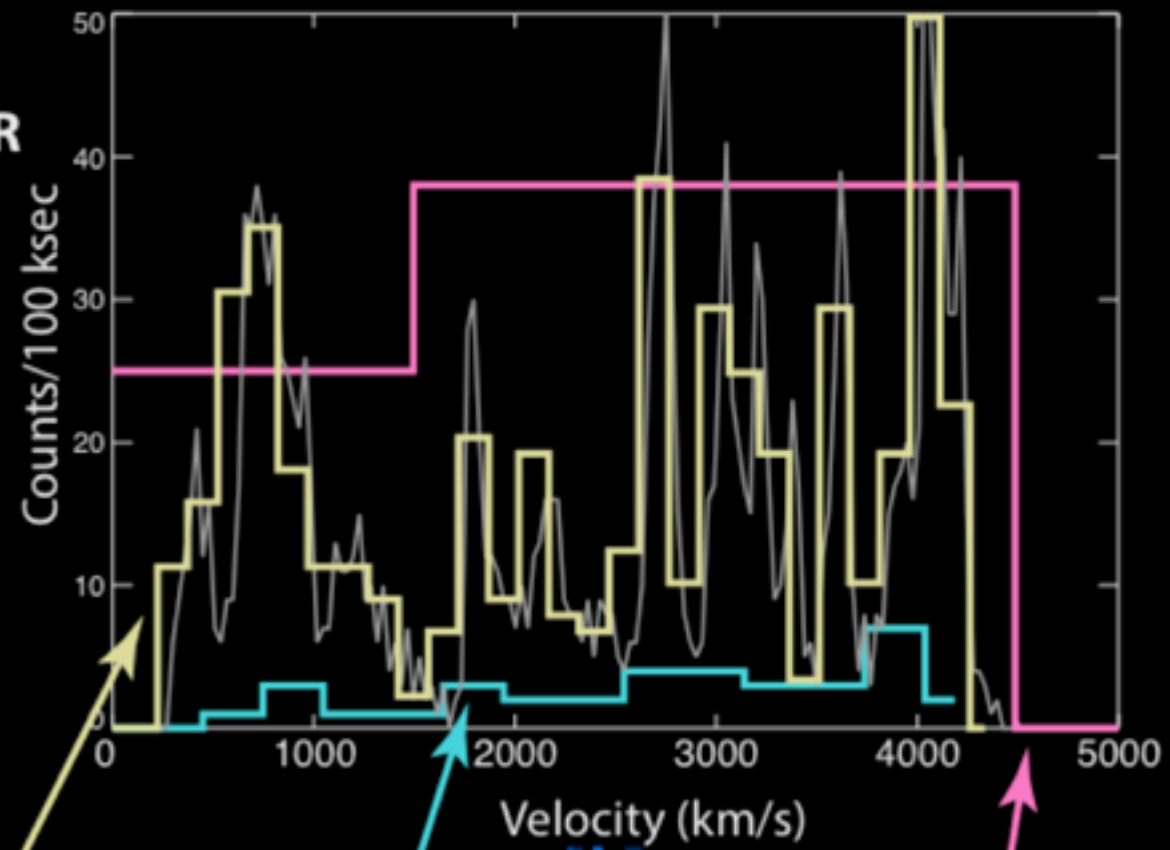
X-ray: NASA/CXC/Rutgers/G.Cassam-Chenaï,
J.Hughes et al.; Radio: NRAO/AUI/NSF/GBT/VLA/
Dyer, Maddalena & Cornwell; Optical: Middlebury
College/F.Winkler, NOAO/AURA/NSF/CTIO
Schmidt & DSS

SN 1006

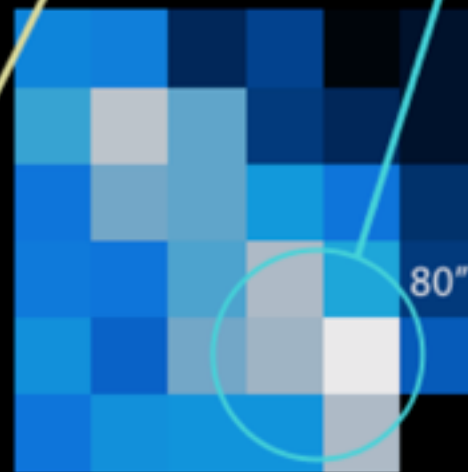
3-D Hydro Simulation Silicon in Tycho-like SNR



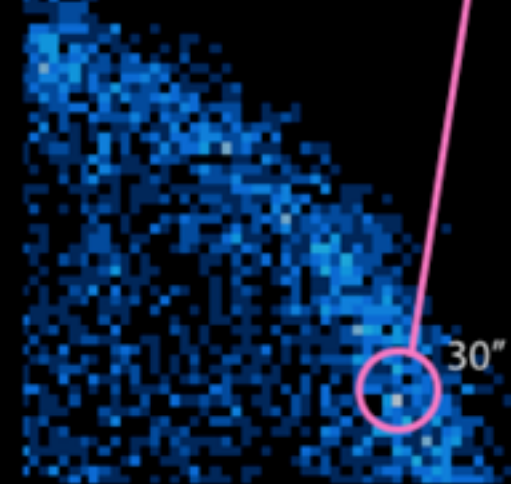
Ferrand et al. (2010)



ATHENA+ XIFU



Astro-H SXS



XMM-Newton EPIC pn

$z = 1.0$

gas temperature

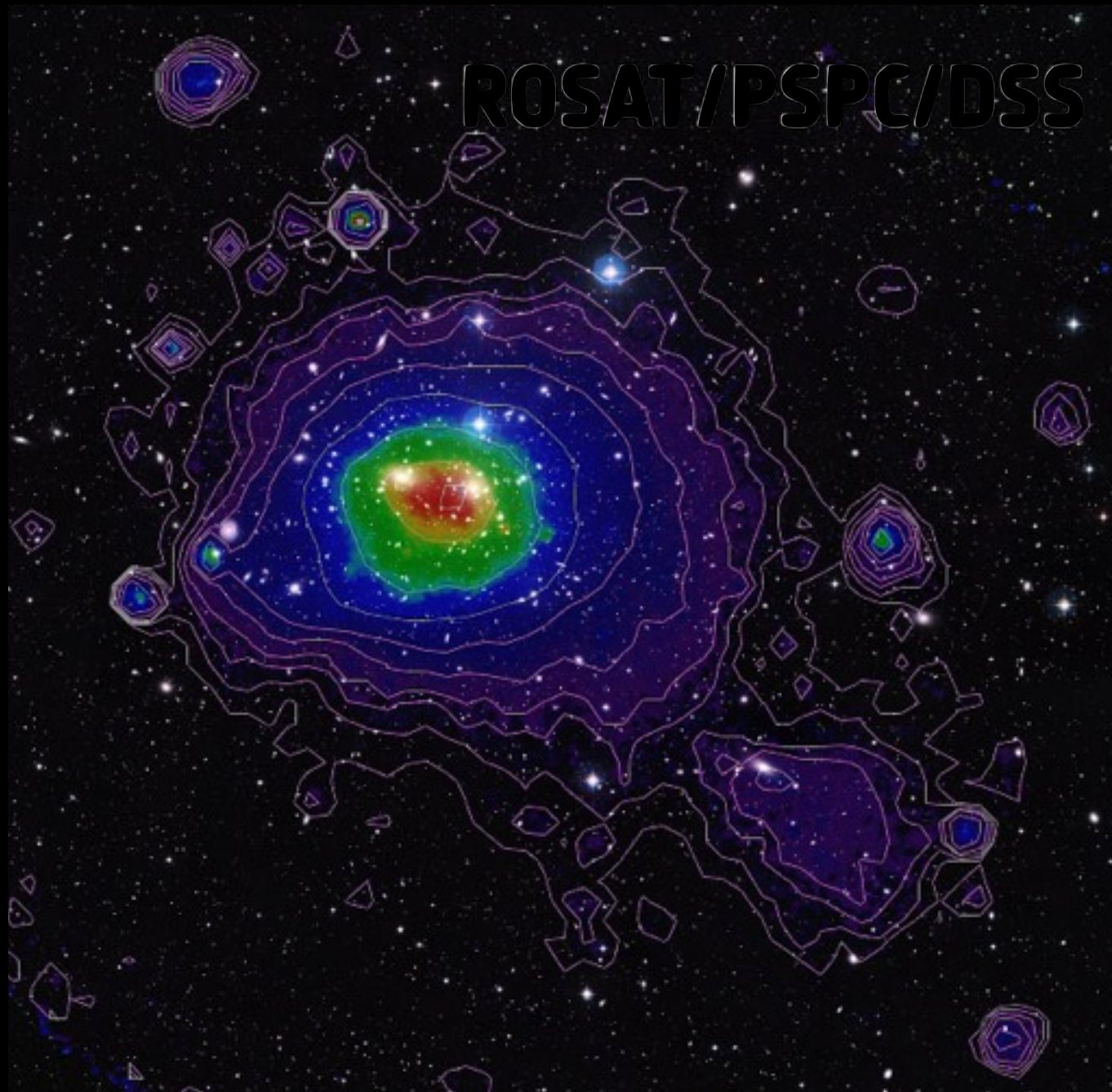
4 Mpc/h

gas density

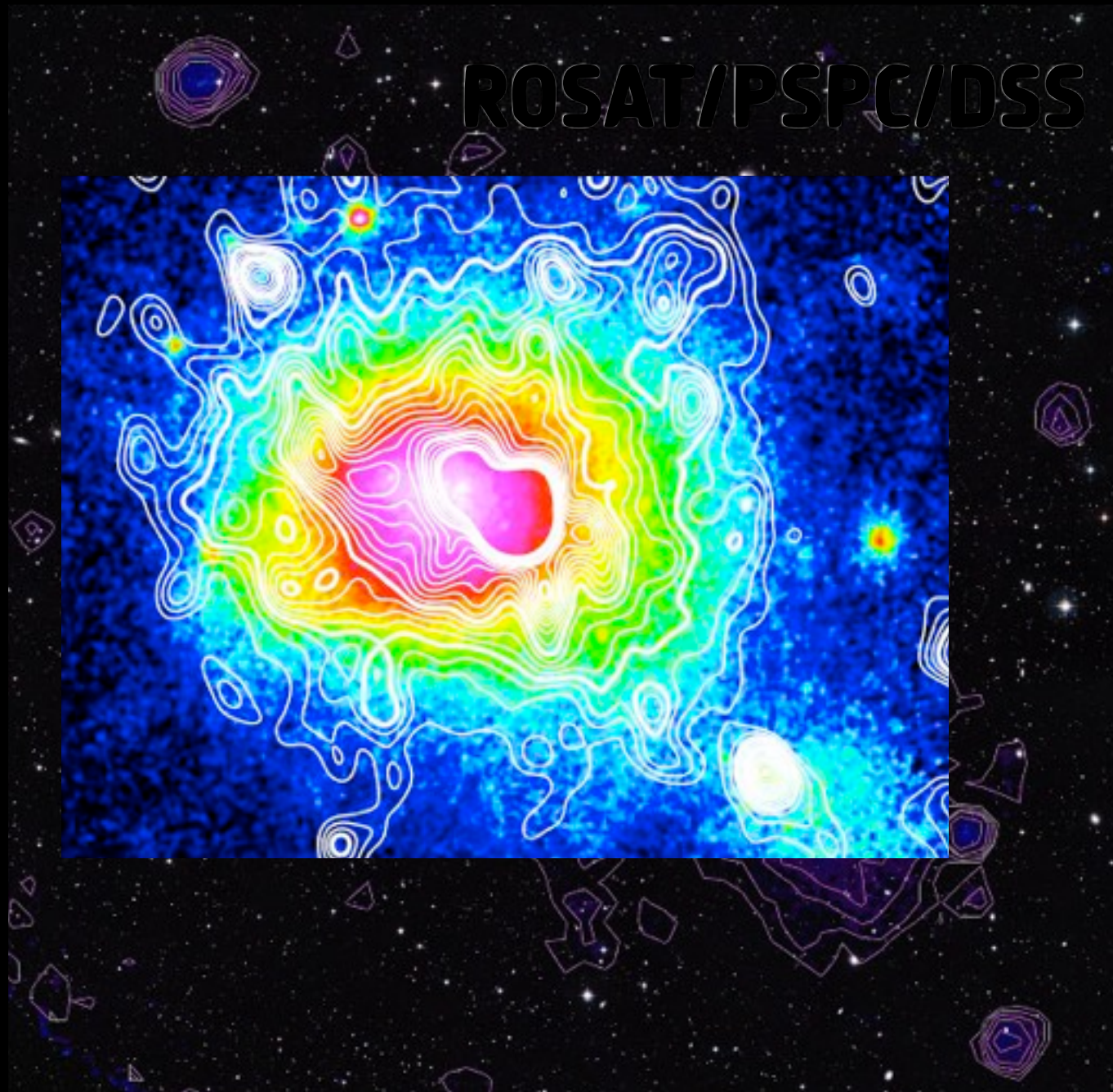
dark matter density

Volker Springel

The halo-turbulence connection

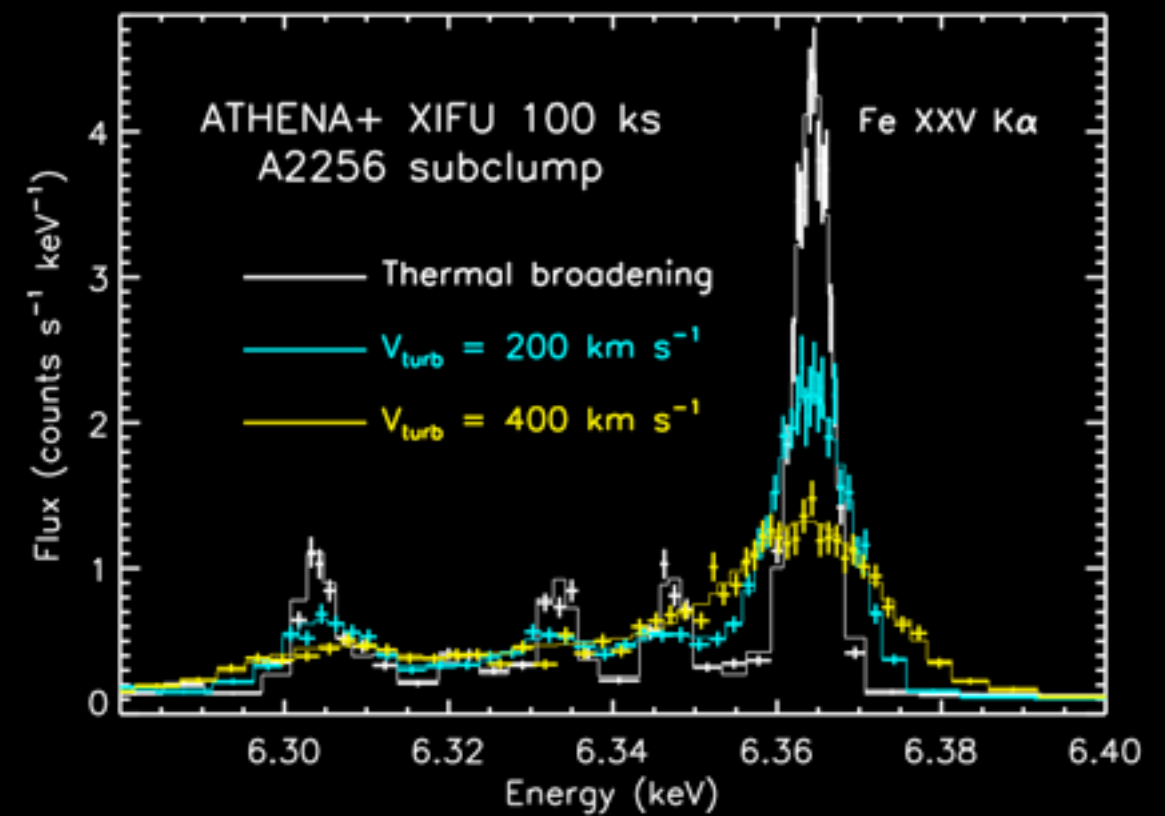
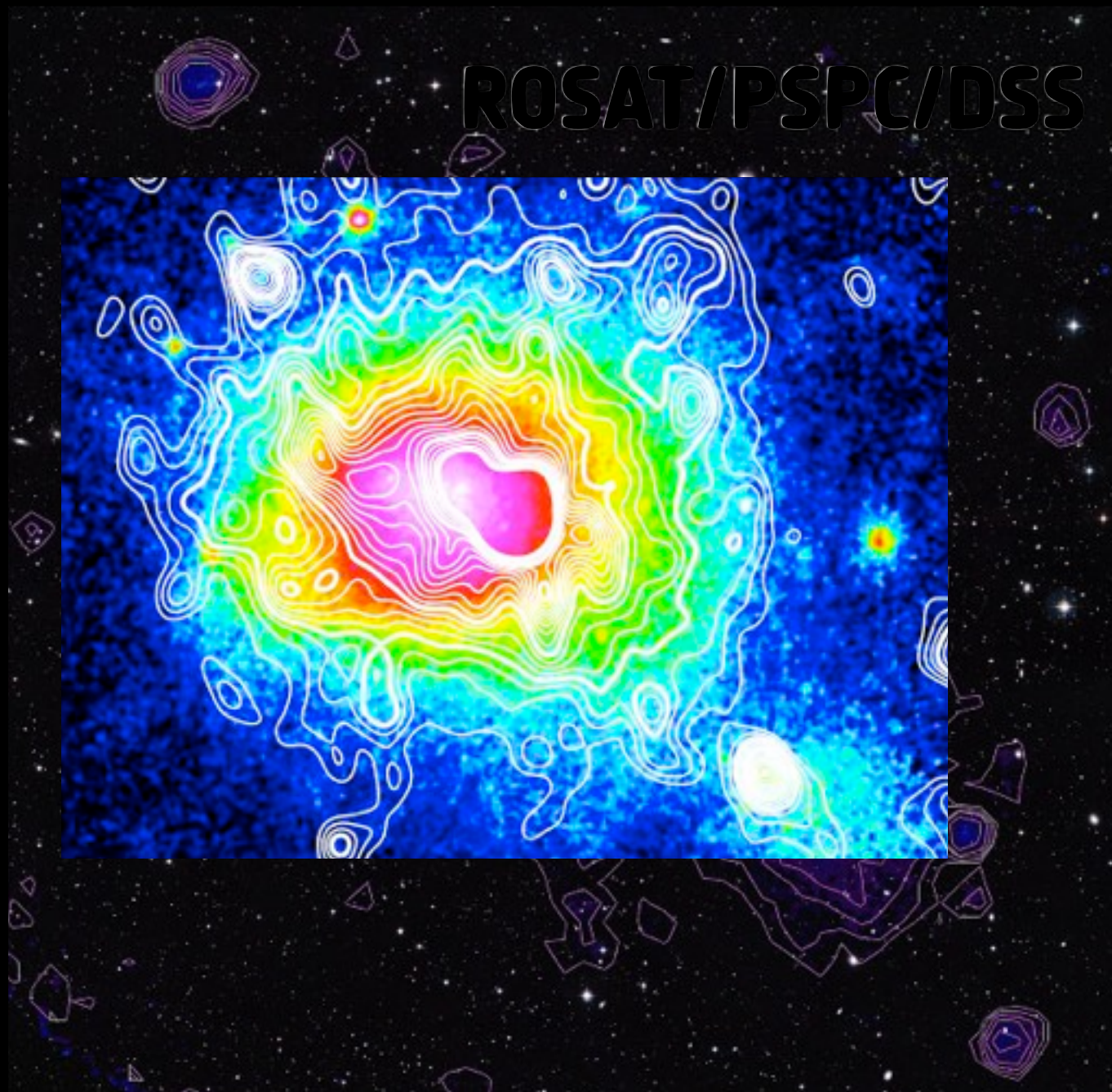


The halo-turbulence connection

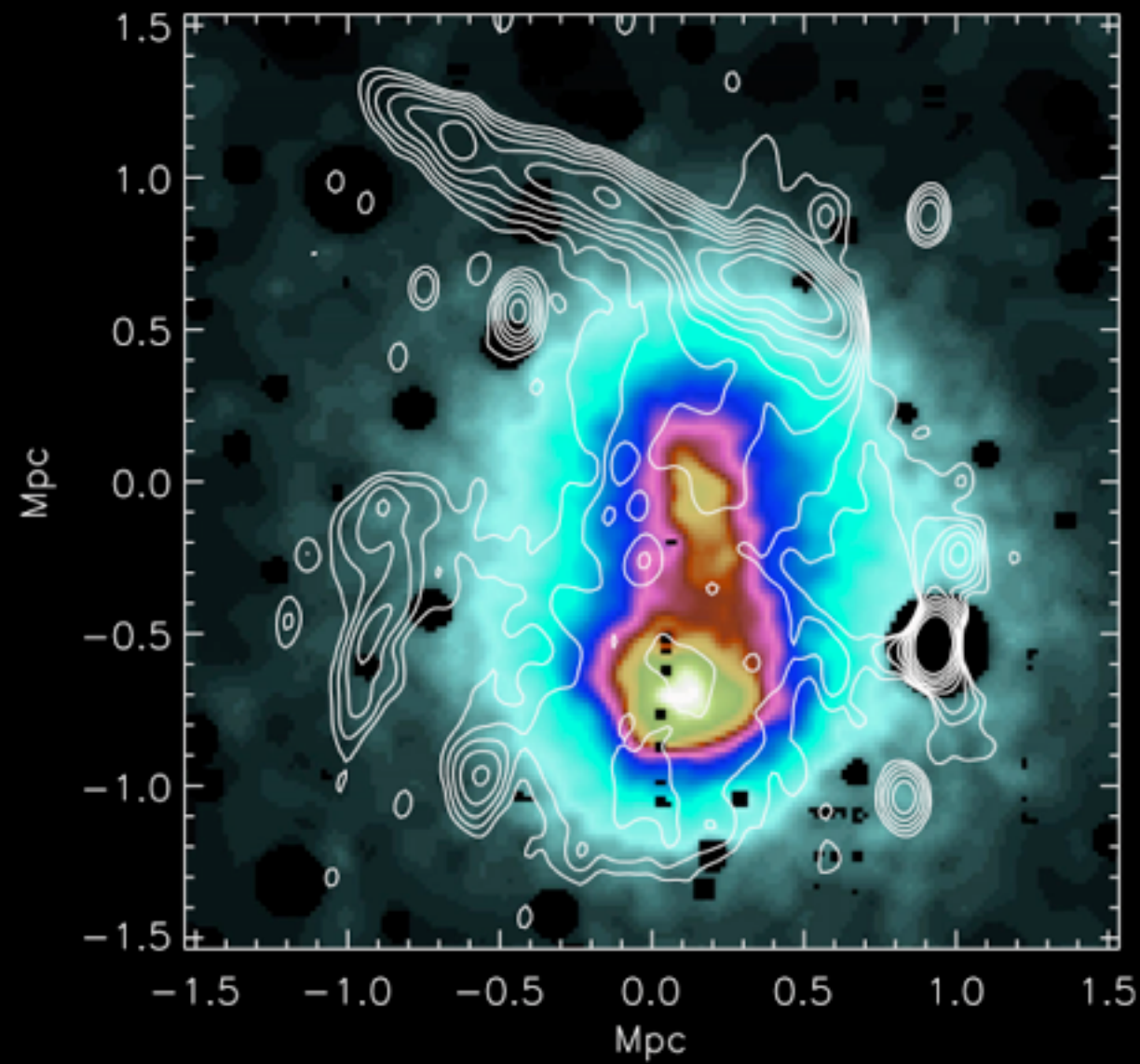


Brown & Rudnick 2010

The halo-turbulence connection

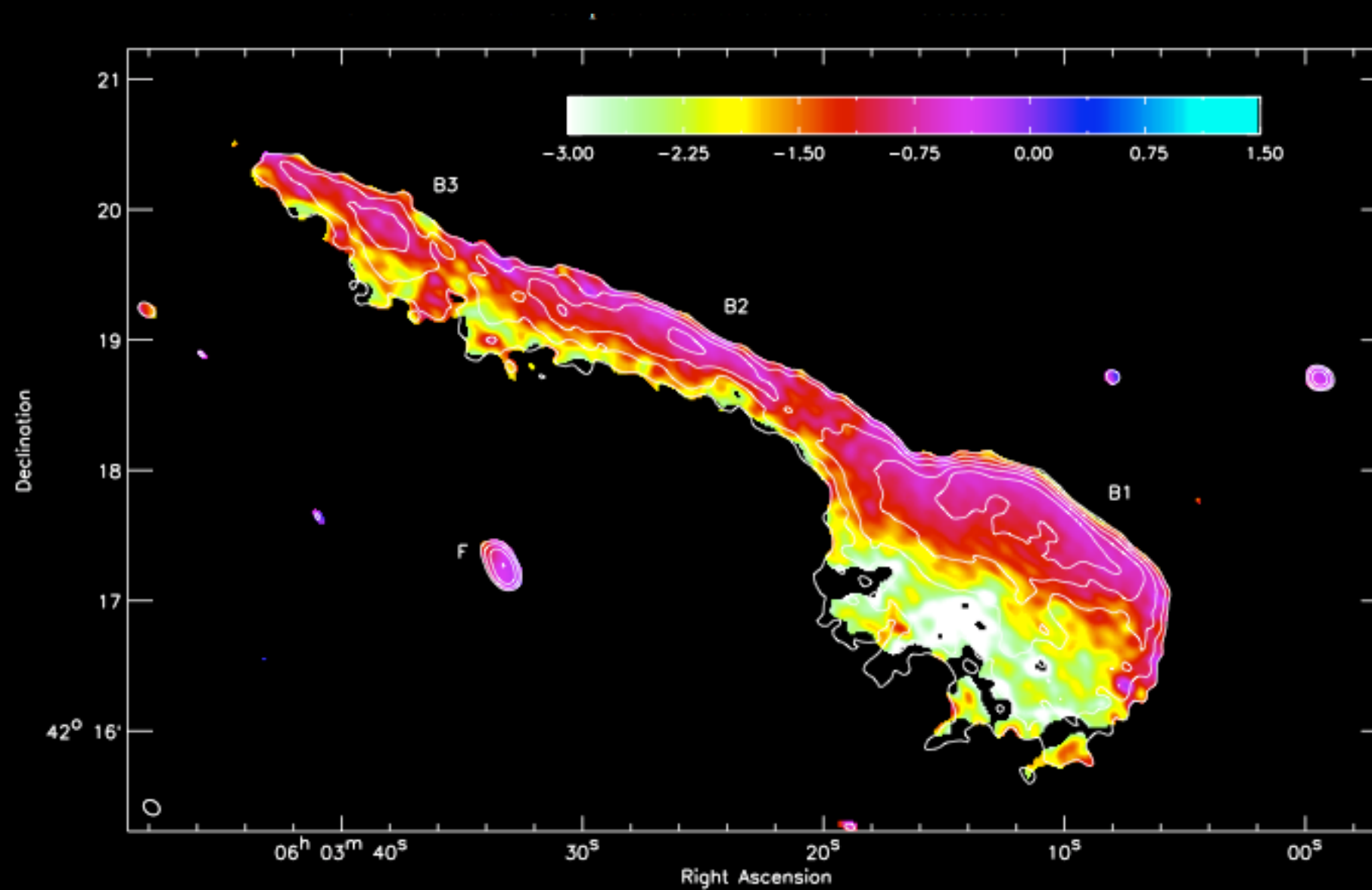


Shock acceleration in relics



Ogrean et al 2013

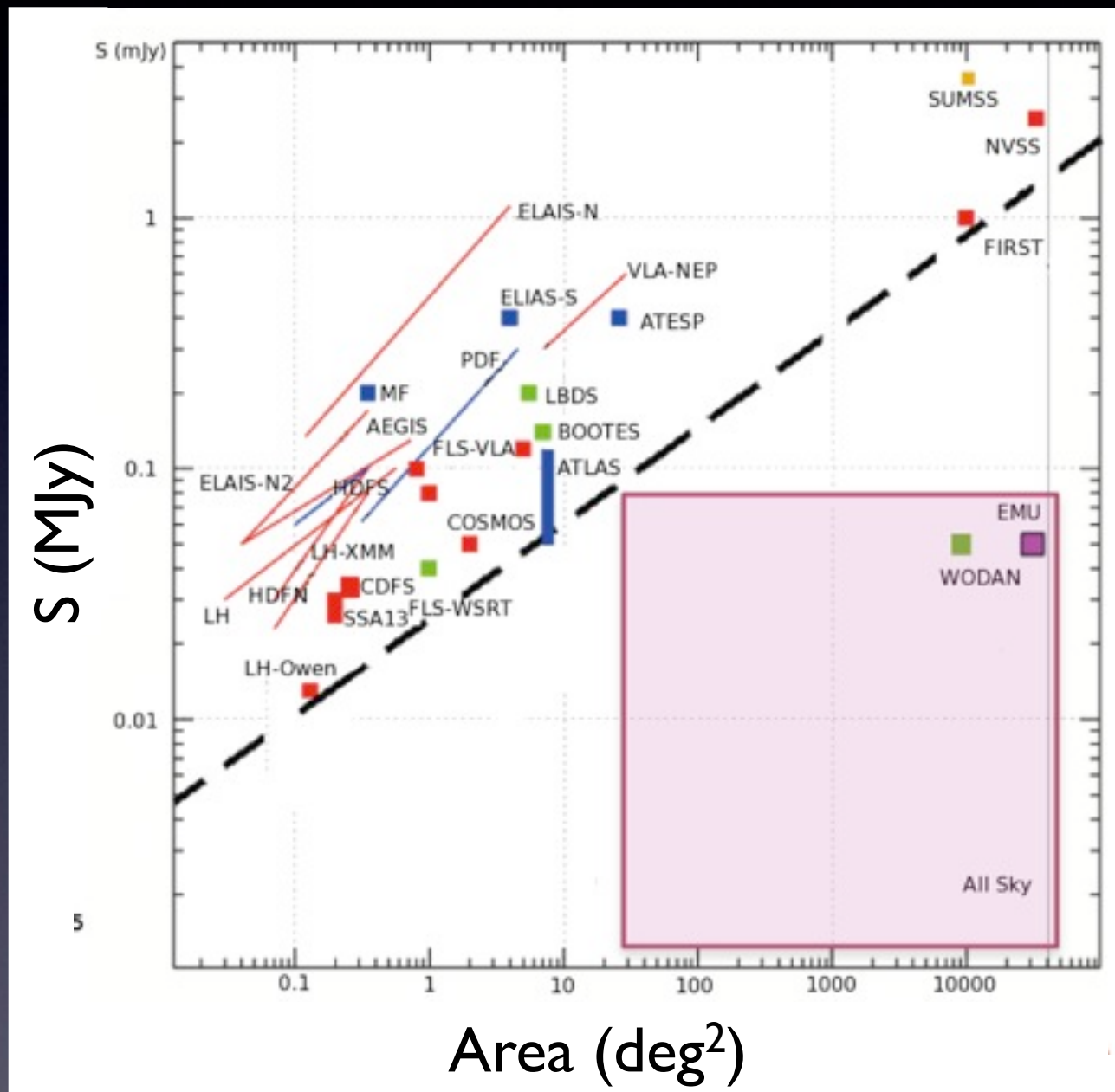
Shock acceleration in relics



van Weeren et al 2012 (WSRT/GMRT)

Some upcoming radio surveys

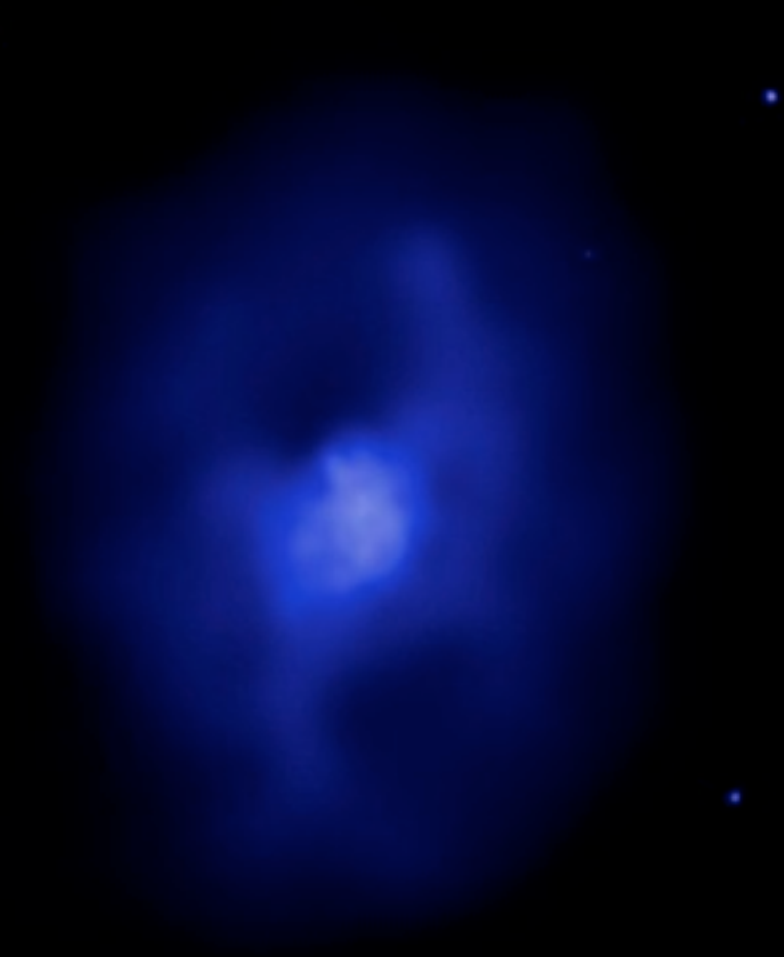
Population studies



Adapted from Norris et al 2012

Survey	Details	Cluster Science
ASKAP-EMU	Southern continuum survey, 1.4GHz, 10'', 10μJy	Haloes, relics, tailed galaxies
WODAN	Northern continuum survey, 1.4GHz, 6'', 10μJy	Haloes, relics, tailed galaxies
LOFAR	Northern sky, 30-240 MHz, 10''	Haloes, relics
MWA	Southern sky, 80-300 MHz, 2.7', 8-9 mJy	Haloes, relics

AGN-ICM interaction



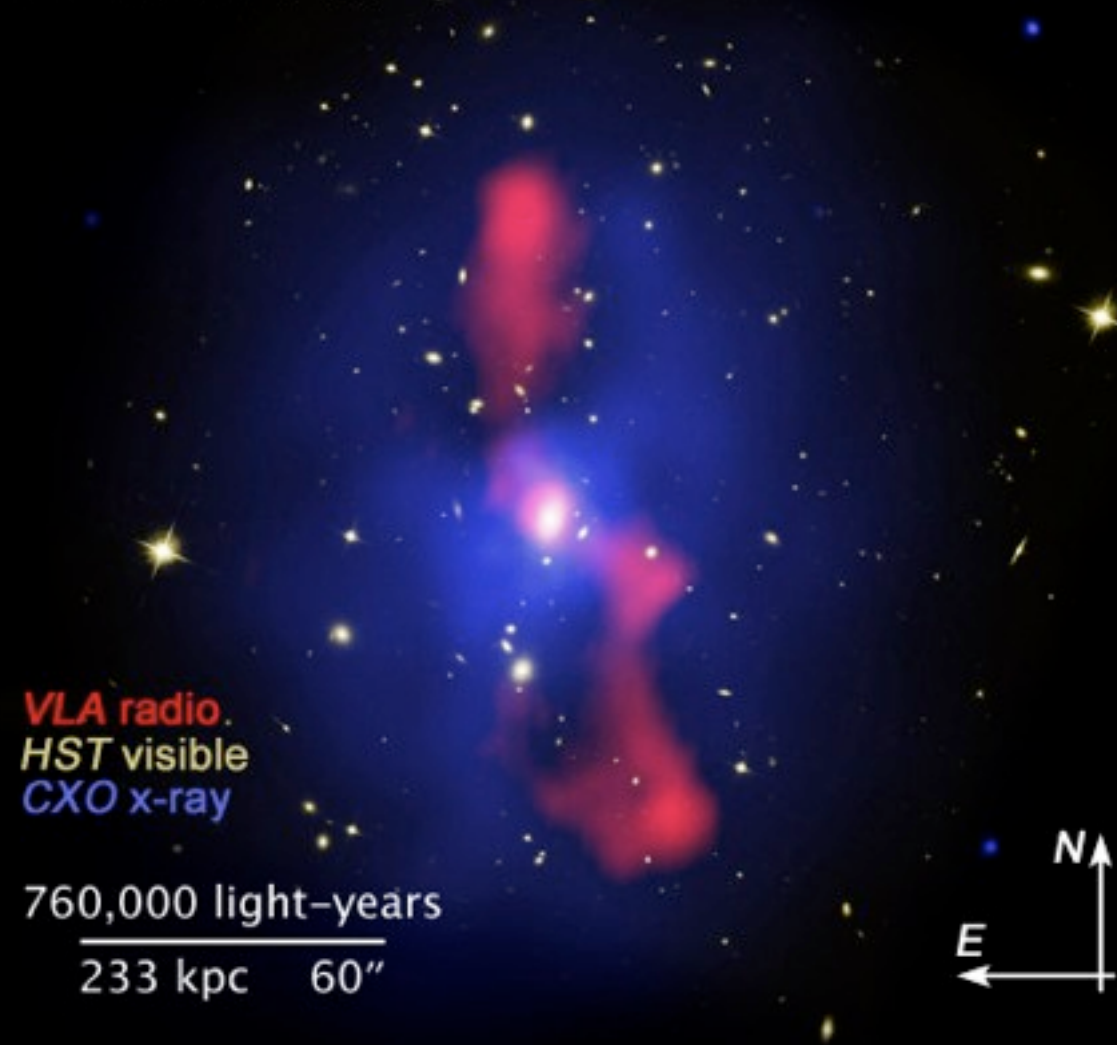
MS0735.6+7421 a $z=0.22$;
NASA, ESA, and B. McNamara



M84 a $z=0.0034$
C. Jones

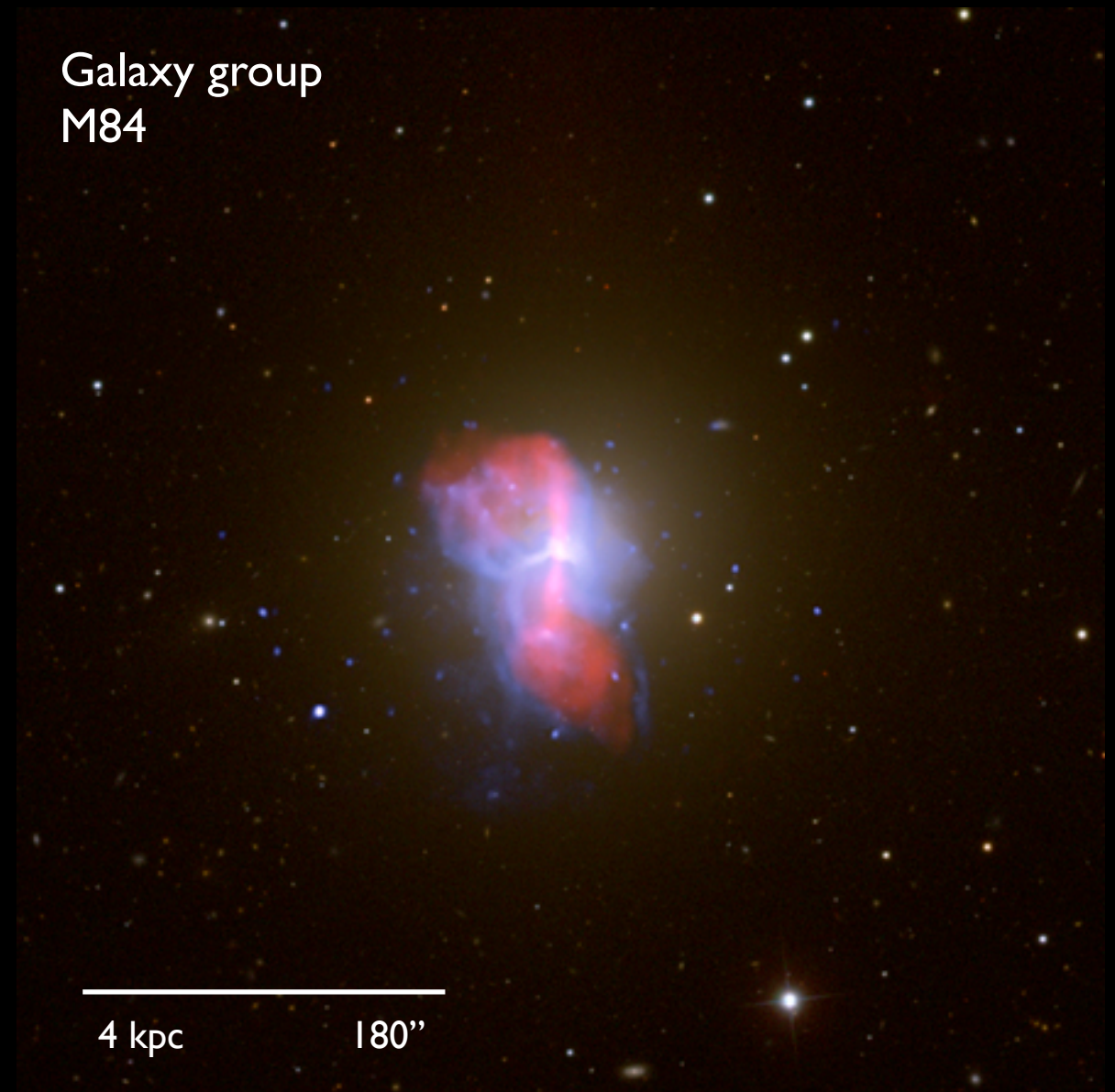
AGN-ICM interaction

Galaxy Cluster
MS 0735.6+7421



MS0735.6+7421 a $z=0.22$;
NASA, ESA, and B. McNamara

Galaxy group
M84

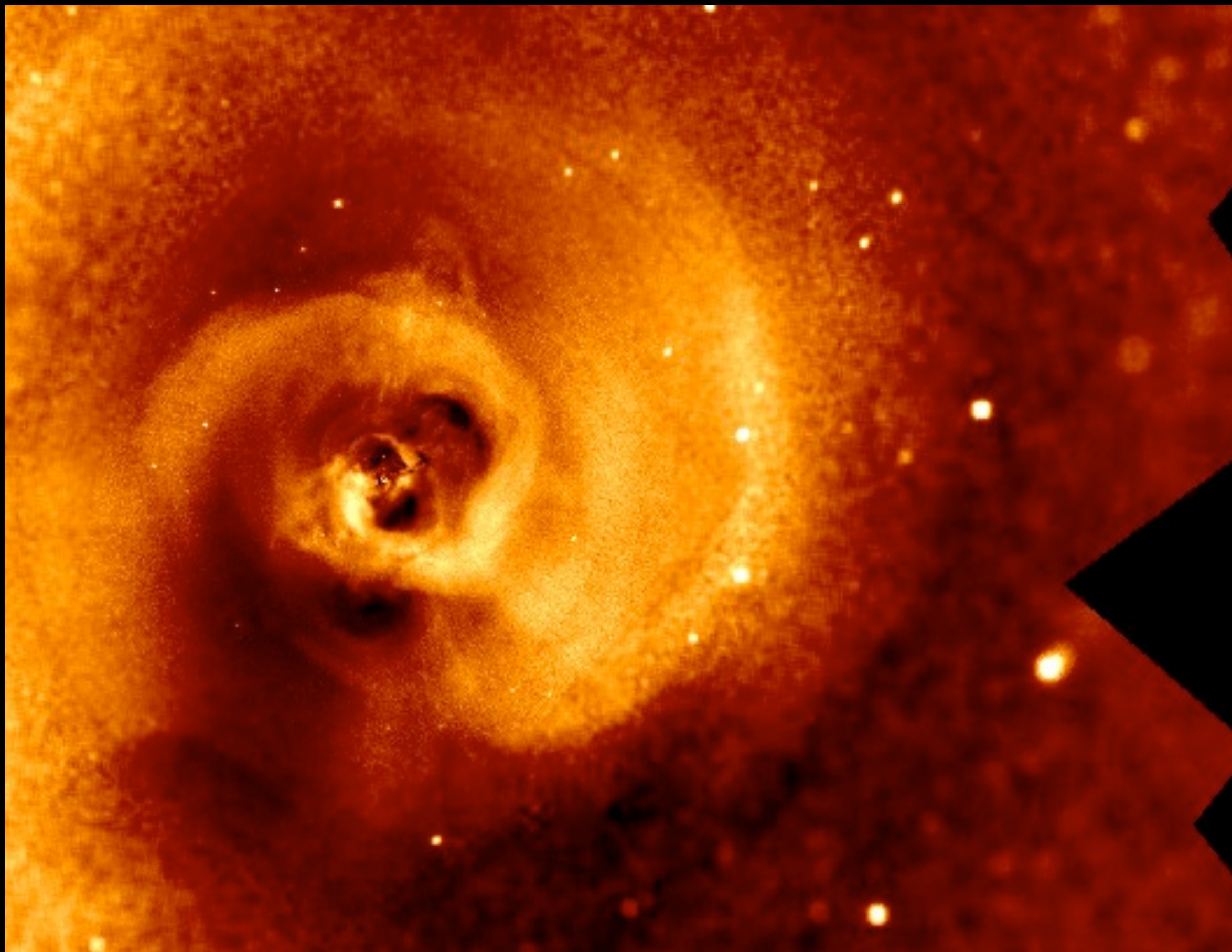


M84 a $z=0.0034$
C. Jones

Conclusions

- ▶ **Athena represents a quantum leap forward in X-ray astronomy**
 - ▶ Unprecedented combination of spatial and spectral resolution
- ▶ **Strong physical connections between radio and X-ray emission**
 - ▶ For many different astrophysical sources
- ▶ **Many synergies between Athena and SKA and pathfinders**
- ▶ **L2 CFP out; proposal due 15 April 2014**
- ▶ **Launch 2028**





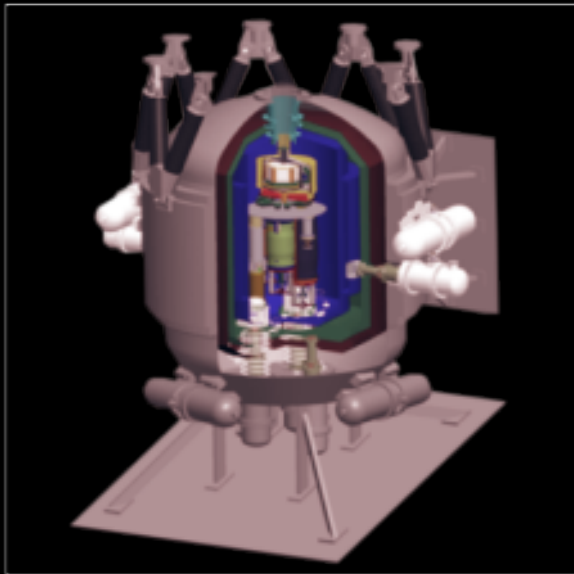
The Advanced Telescope for High ENergy Astrophysics

L2 orbit Ariane V

Mass < 5100 kg

Power 2500 W

5 year mission

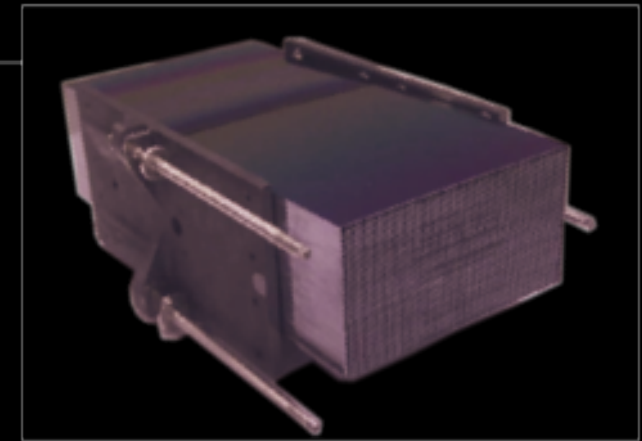
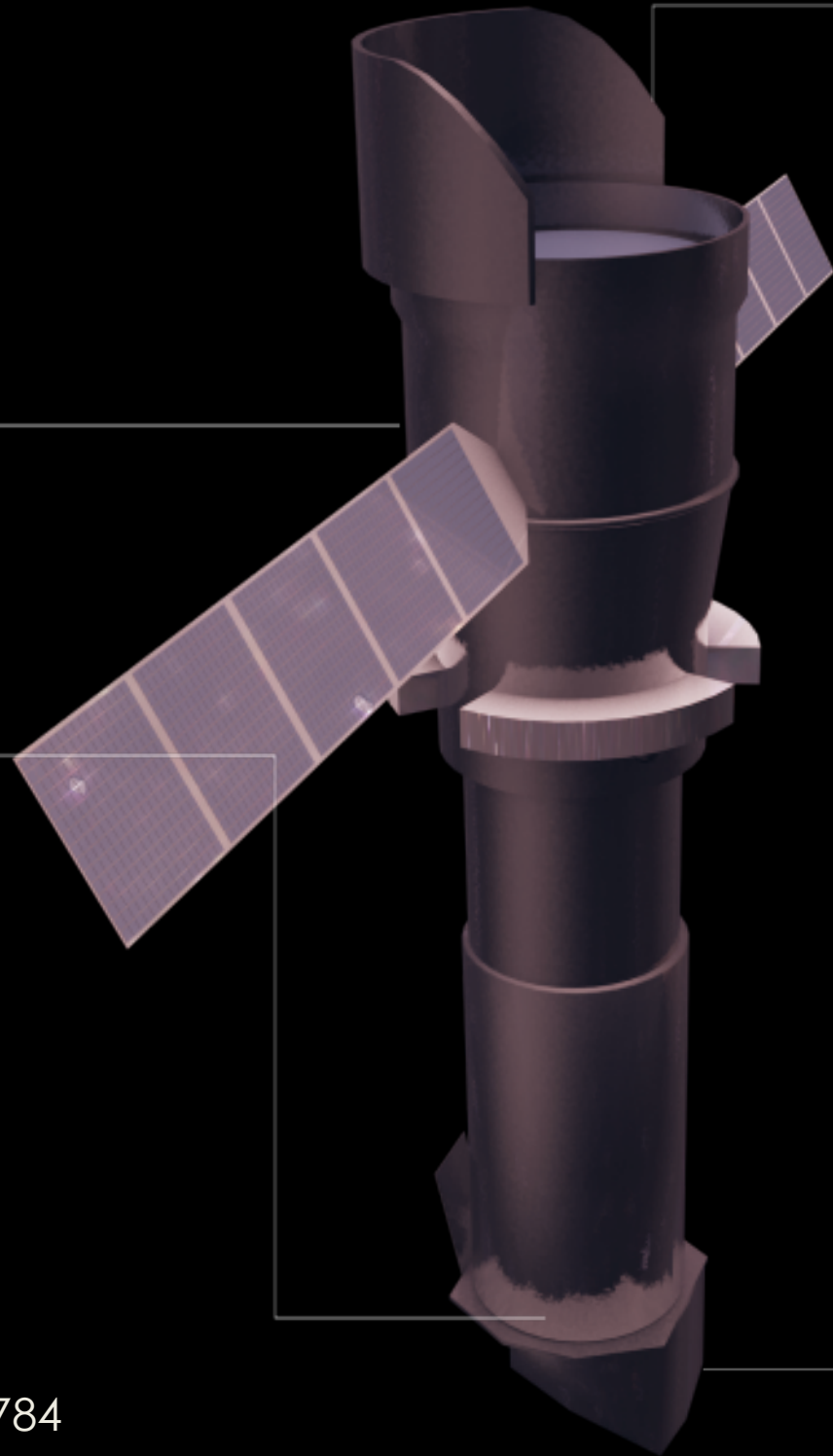


X-ray Integral Field Unit:

ΔE : 2.5 eV

Field of View: 5 arcmin

Operating temp: 50 mk



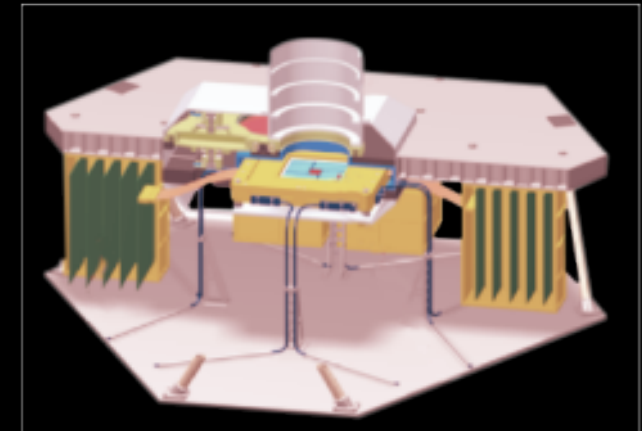
Silicon Pore Optics:

2 m² at 1 keV

5 arcsec HEW

Focal length: 12 m

Sensitivity: $3 \cdot 10^{-17}$ erg cm⁻² s⁻¹



Wide Field Imager:

ΔE : 125 eV

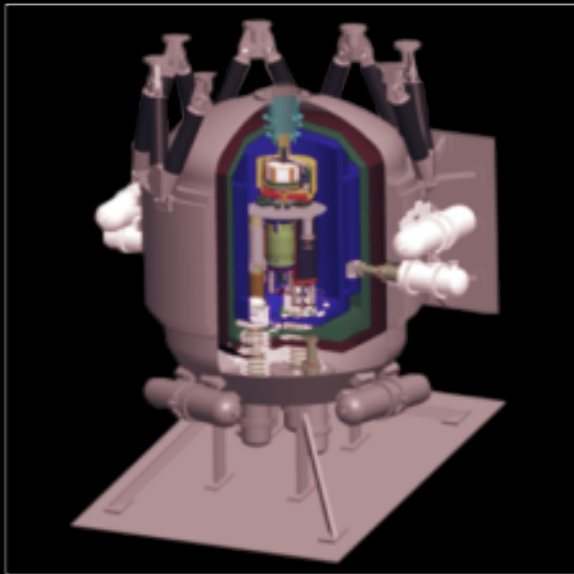
Field of View: 40 arcmin

High countrate capability

The Advanced Telescope for High ENergy Astrophysics

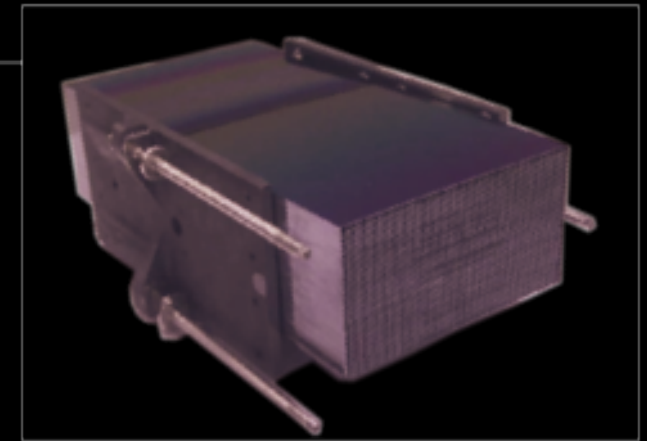
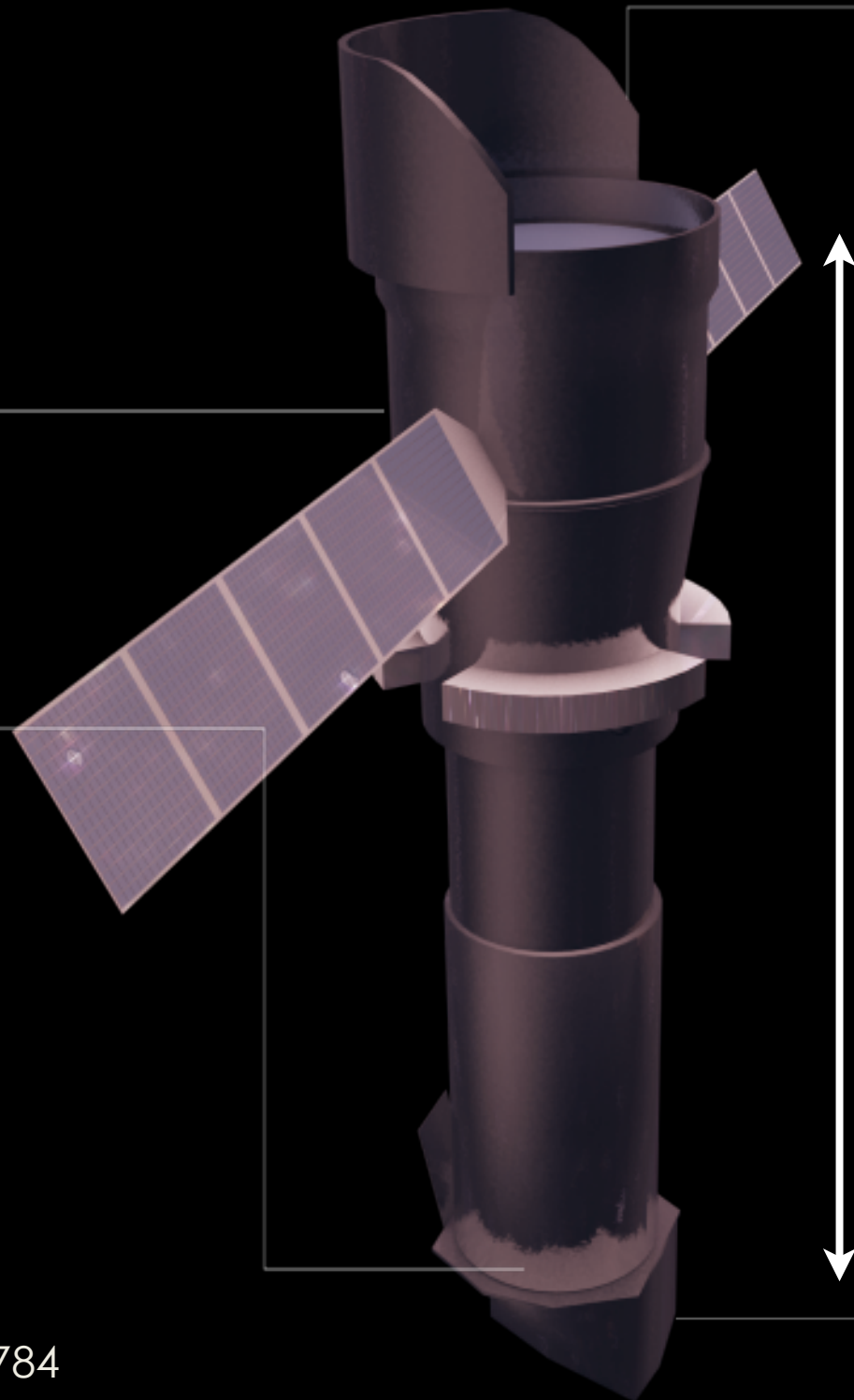
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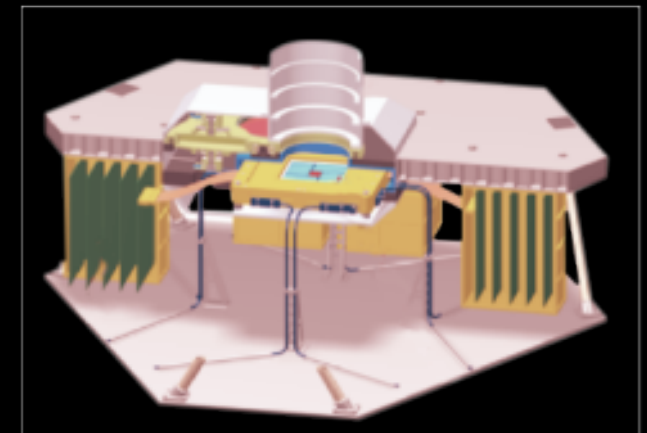
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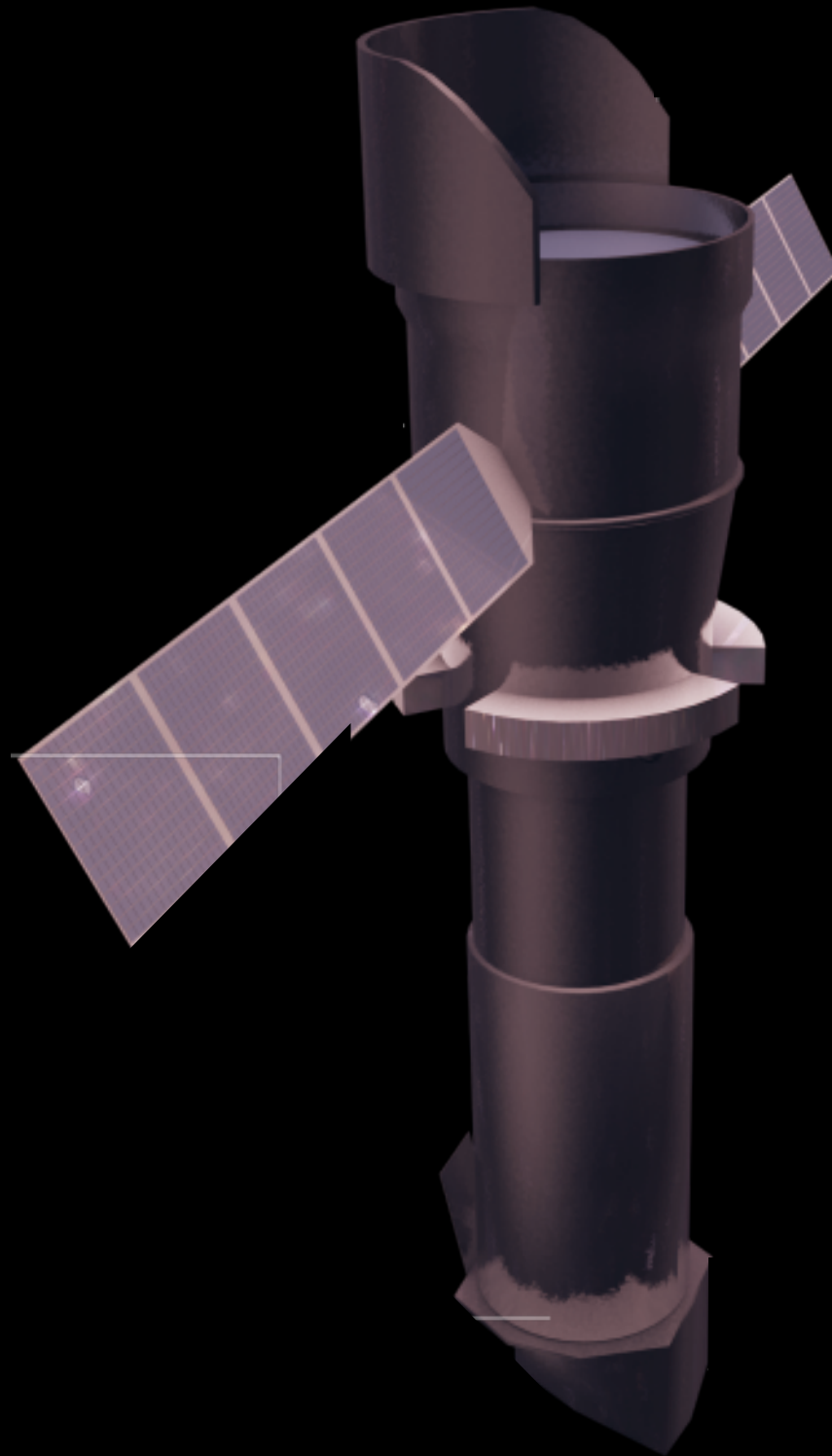
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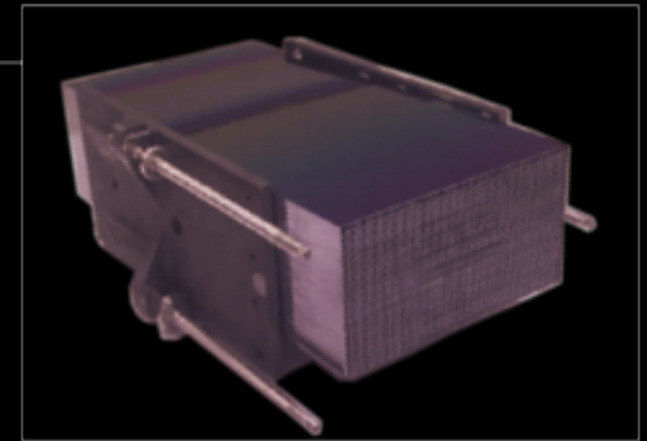
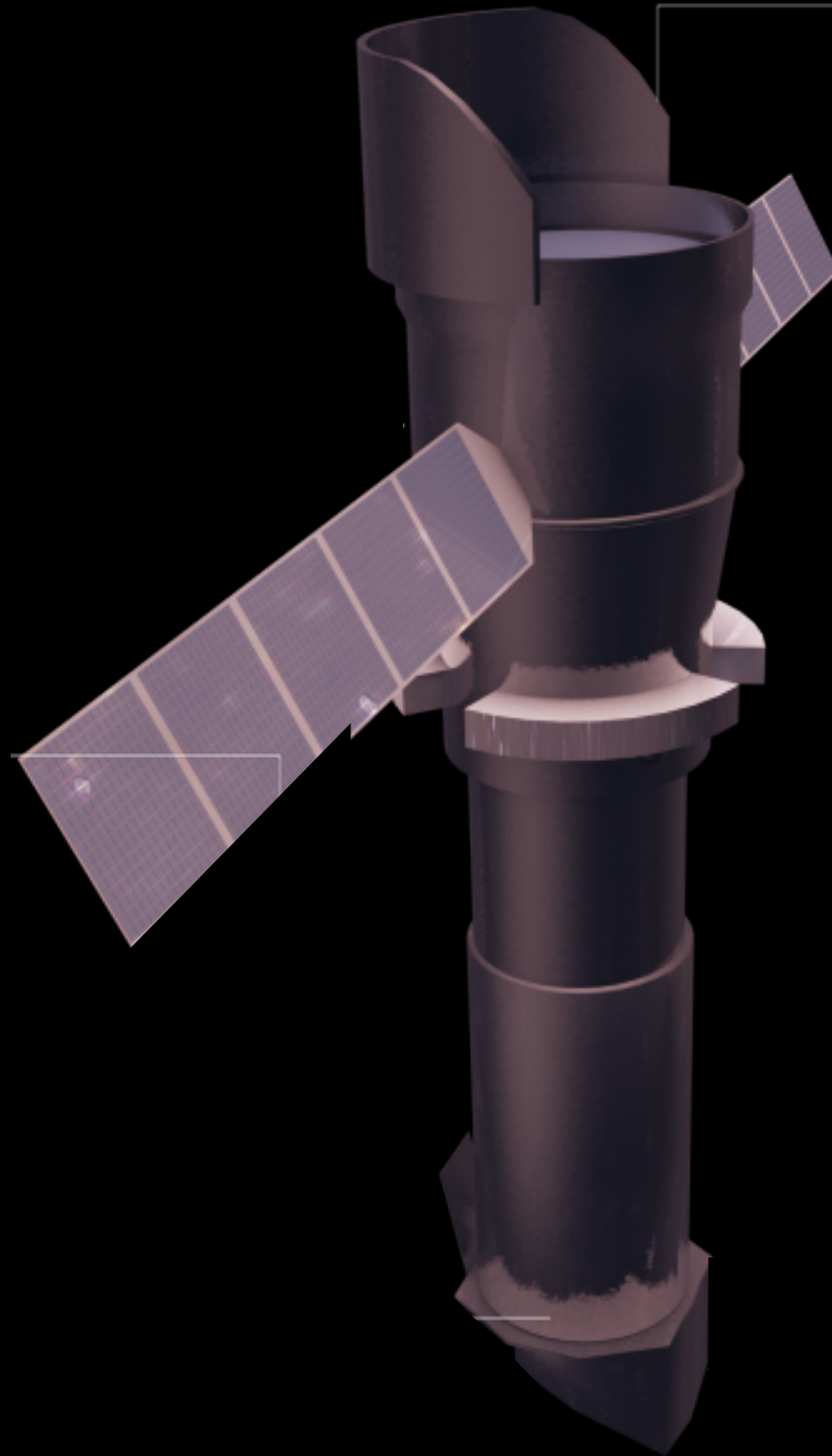
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The Advanced Telescope for High ENergy Astrophysics



The Advanced Telescope for High ENergy Astrophysics



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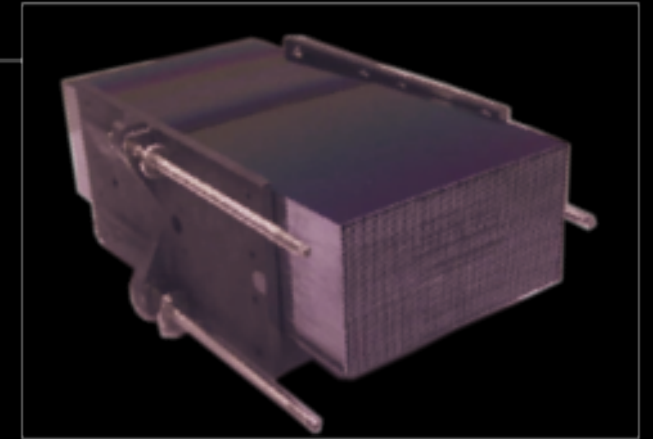
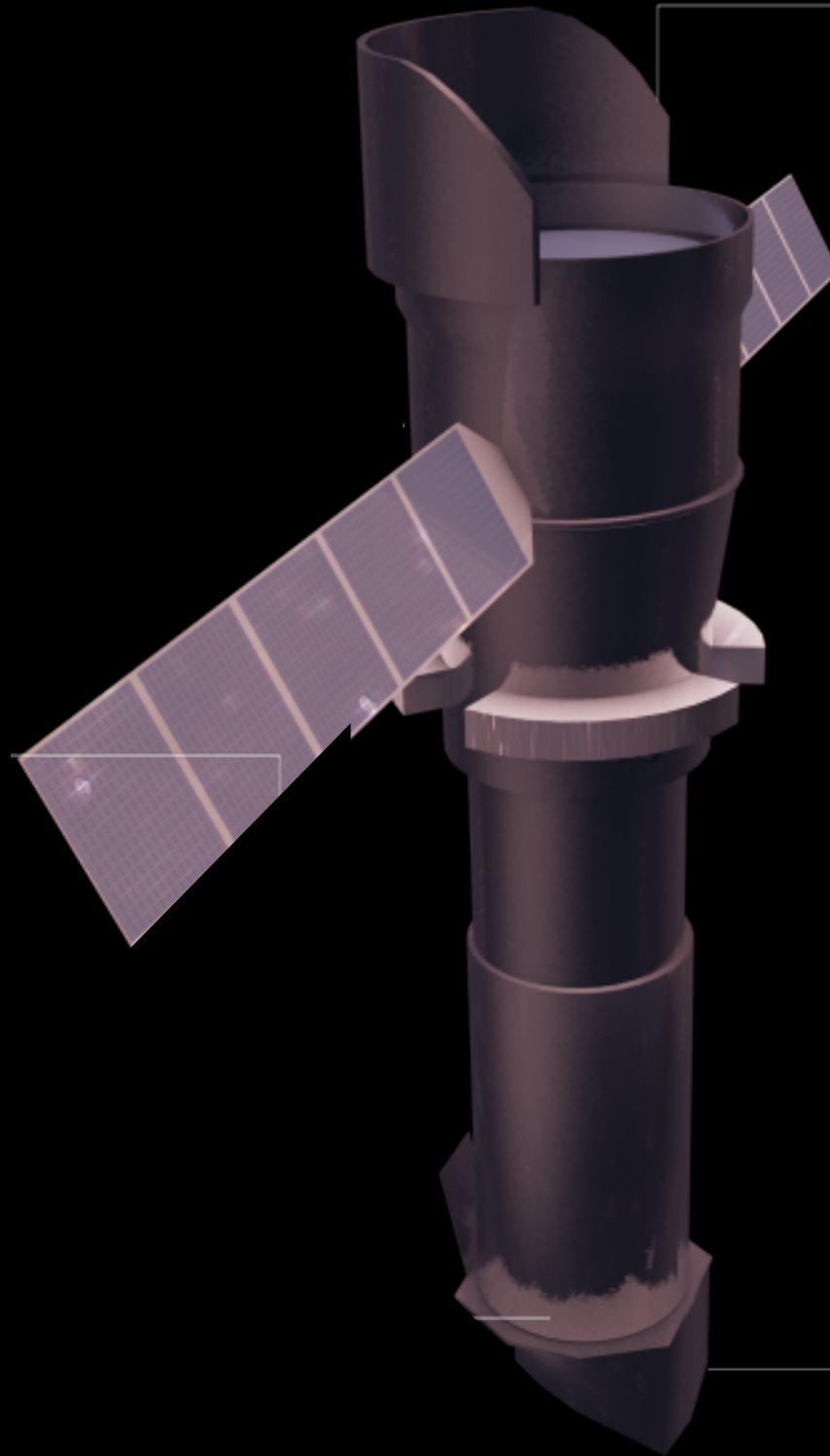
2 m² at 1 keV

5 arcsec HEW

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Sensitivity: 3×10^{-17} erg cm⁻² s⁻¹

The Advanced Telescope for High ENergy Astrophysics



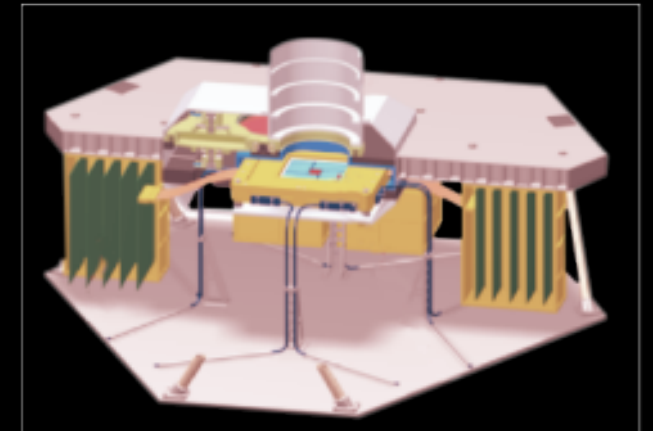
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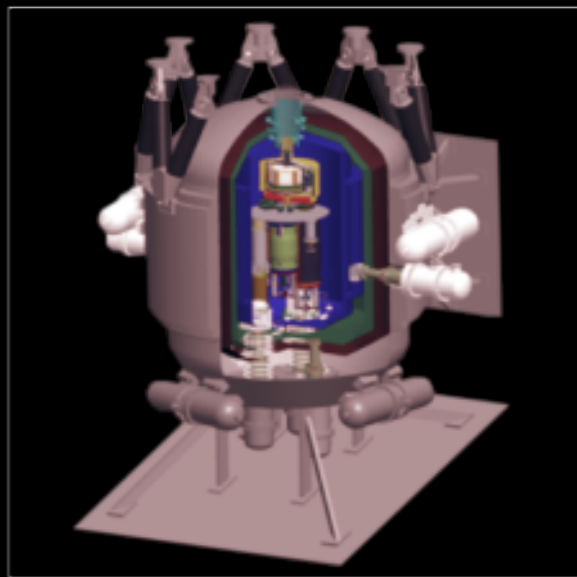
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The Advanced Telescope for High ENergy Astrophysics



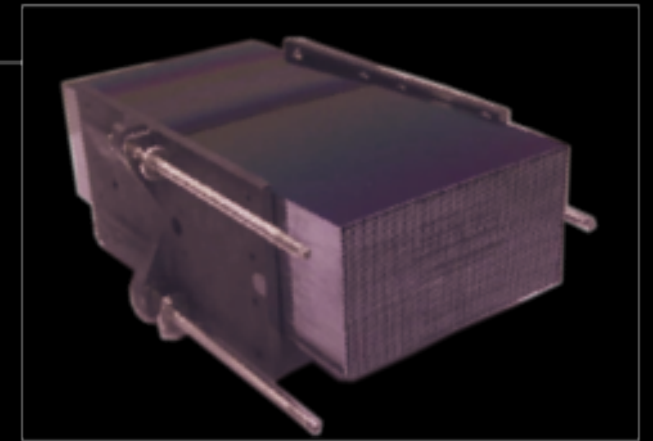
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ΔE : 2.5 eV

Field of View: 5 arcmin

Operating temp: 50 mk

Barret et al., 2013 arXiv:1308.6784



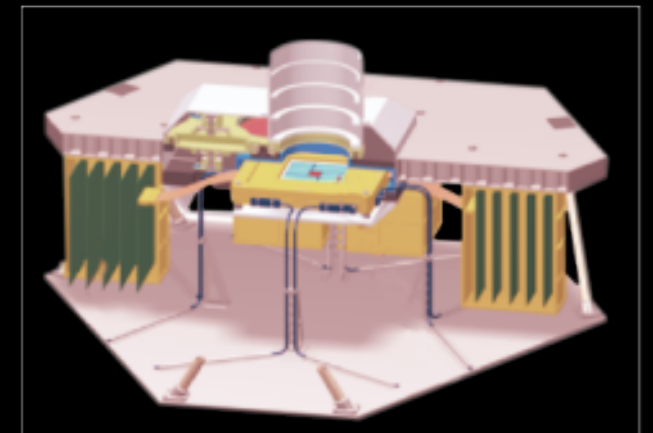
Silicon Pore Optics:

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5 arcsec HEW

Focal length: 12 m

Sensitivity: 3 10⁻¹⁷ erg cm⁻² s⁻¹



Wide Field Imager:

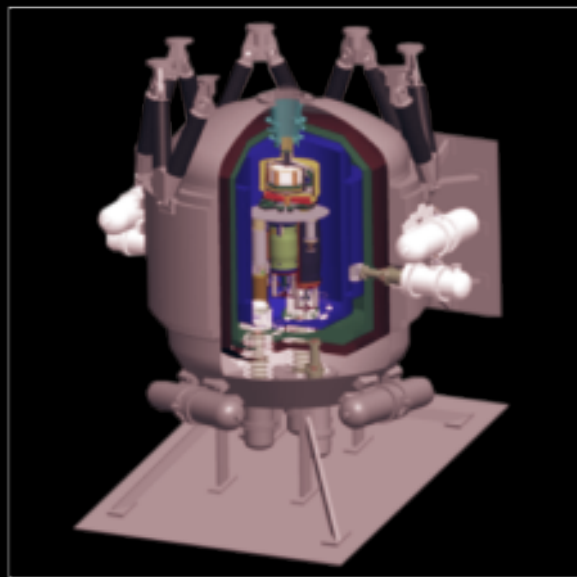
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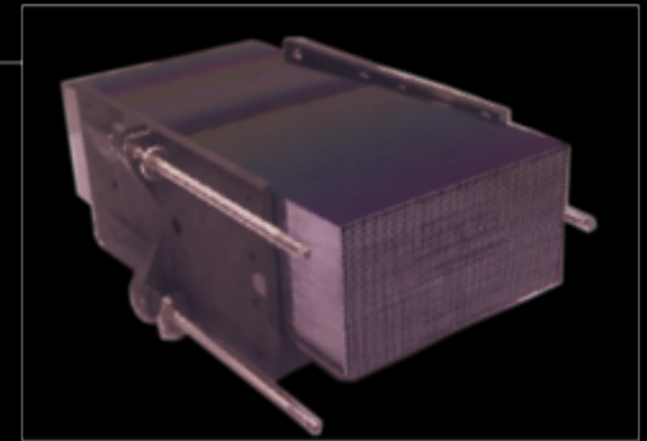
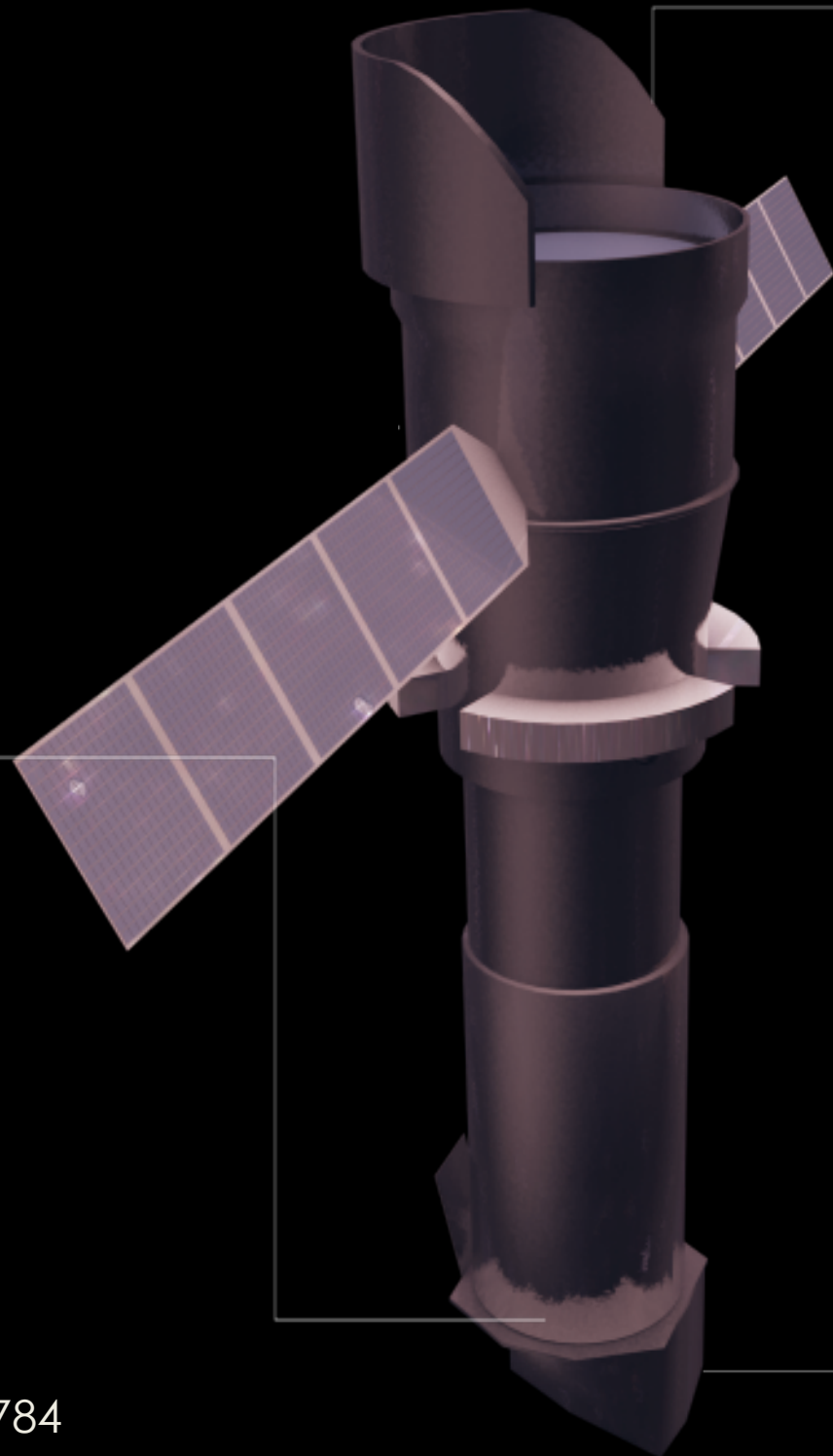
Rau et al. 2013 arXiv1307.1709

The Advanced Telescope for High ENergy Astrophysics



X-ray Integral Field Unit:

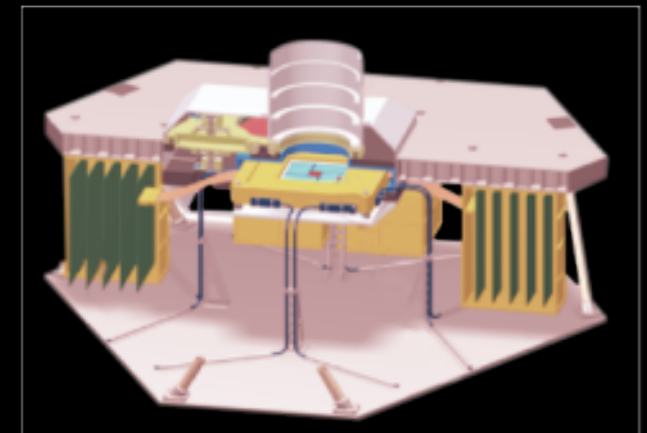
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Sensitivity: $3 \cdot 10^{-17}$ erg cm⁻² s⁻¹

12m



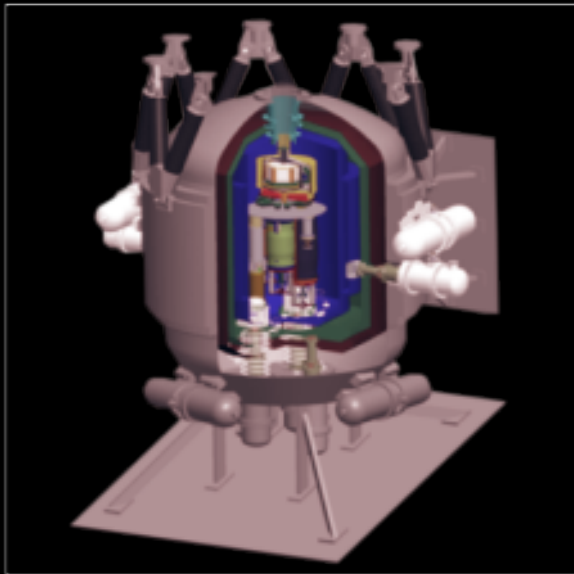
Wide Field Imager:

ΔE : 125 eV
Field of View: 40 arcmin
High countrate capability

The Advanced Telescope for High ENergy Astrophysics

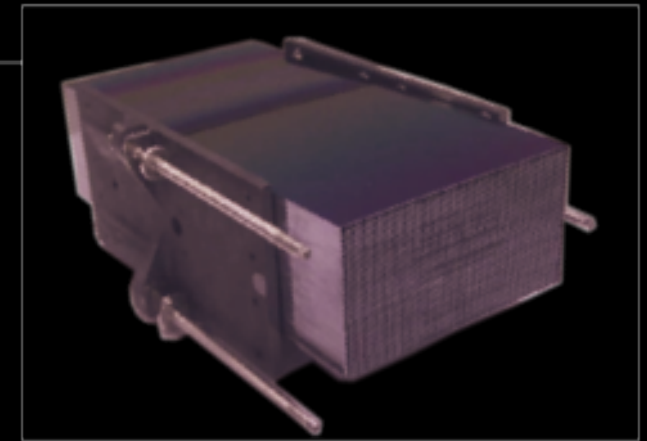
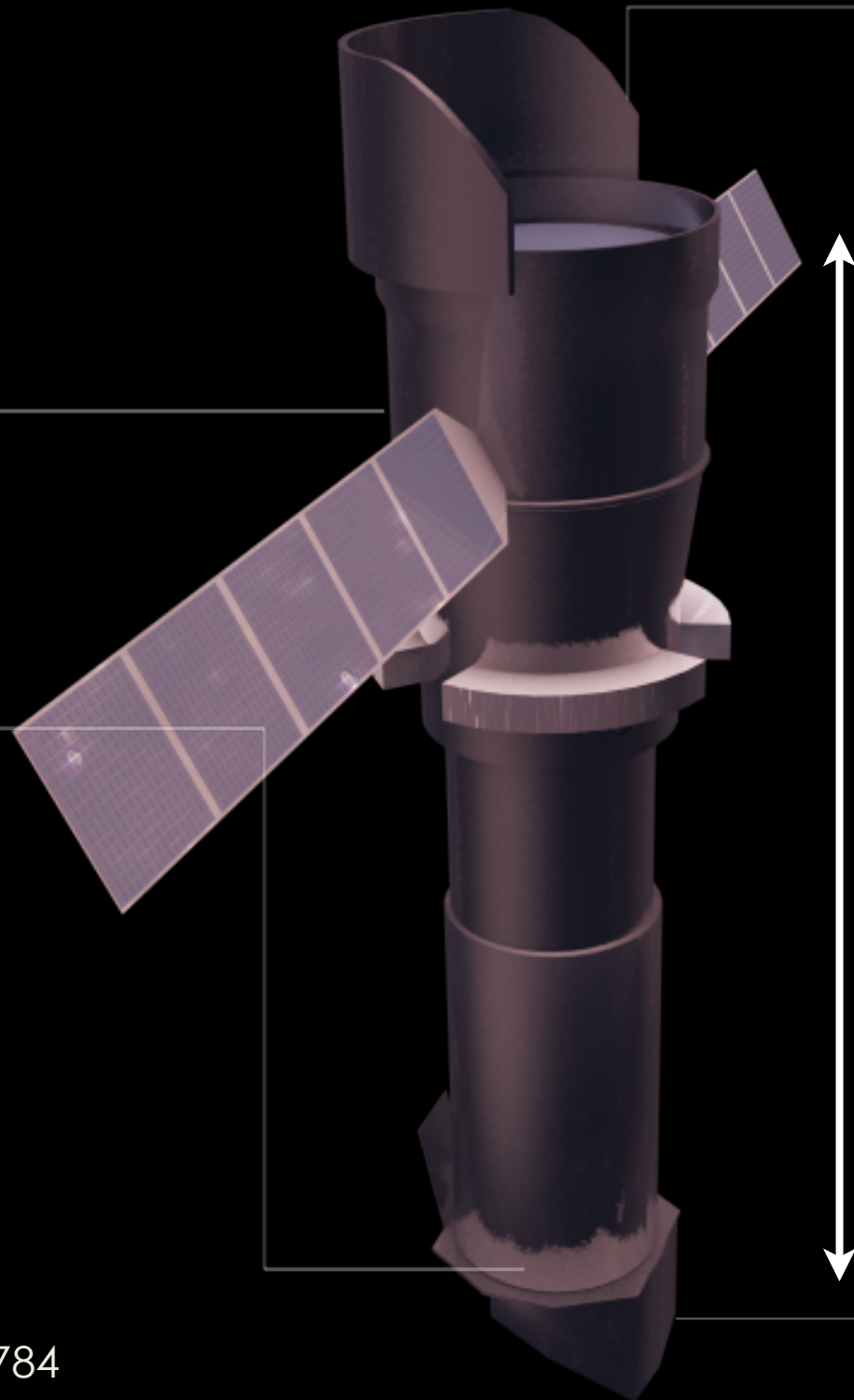
L2 orbit Ariane V

Mass < 5100 kg
Power 2500 W
5 year mission



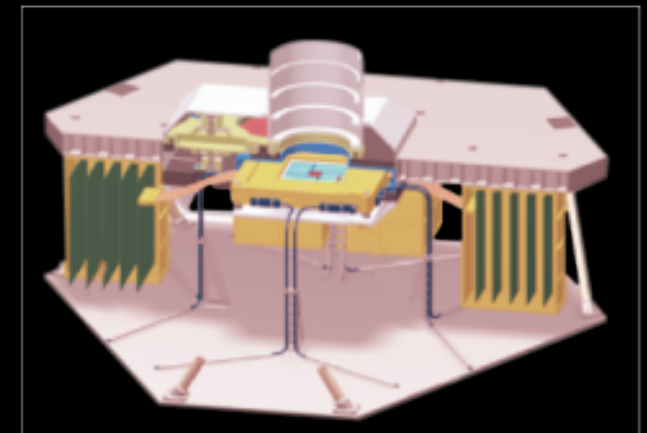
X-ray Integral Field Unit:

ΔE : 2.5 eV
Field of View: 5 arcmin
Operating temp: 50 mk



Silicon Pore Optics:

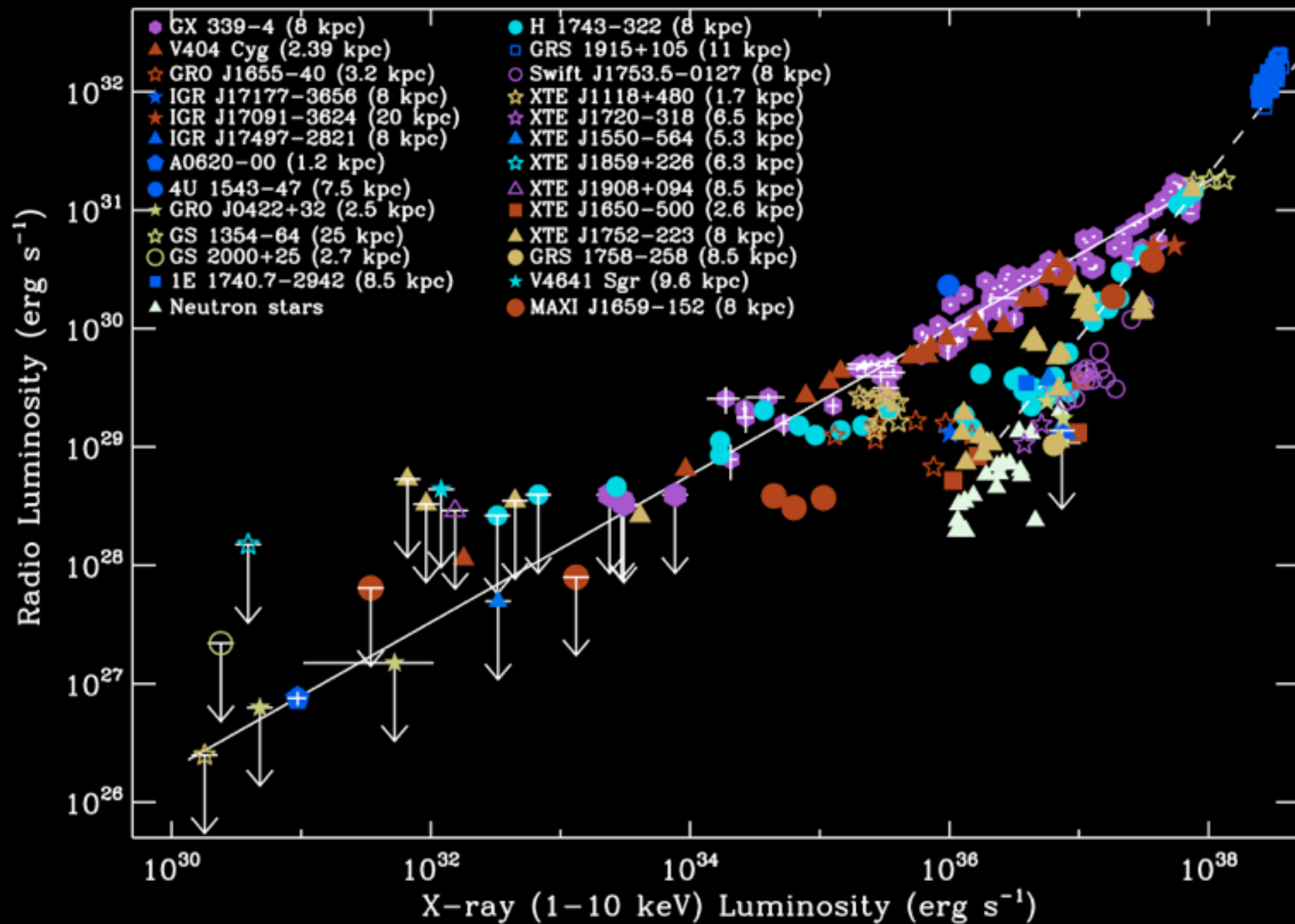
2 m² at 1 keV
5 arcsec HEW
Focal length: 12 m
Sensitivity: $3 \cdot 10^{-17}$ erg cm⁻² s⁻¹



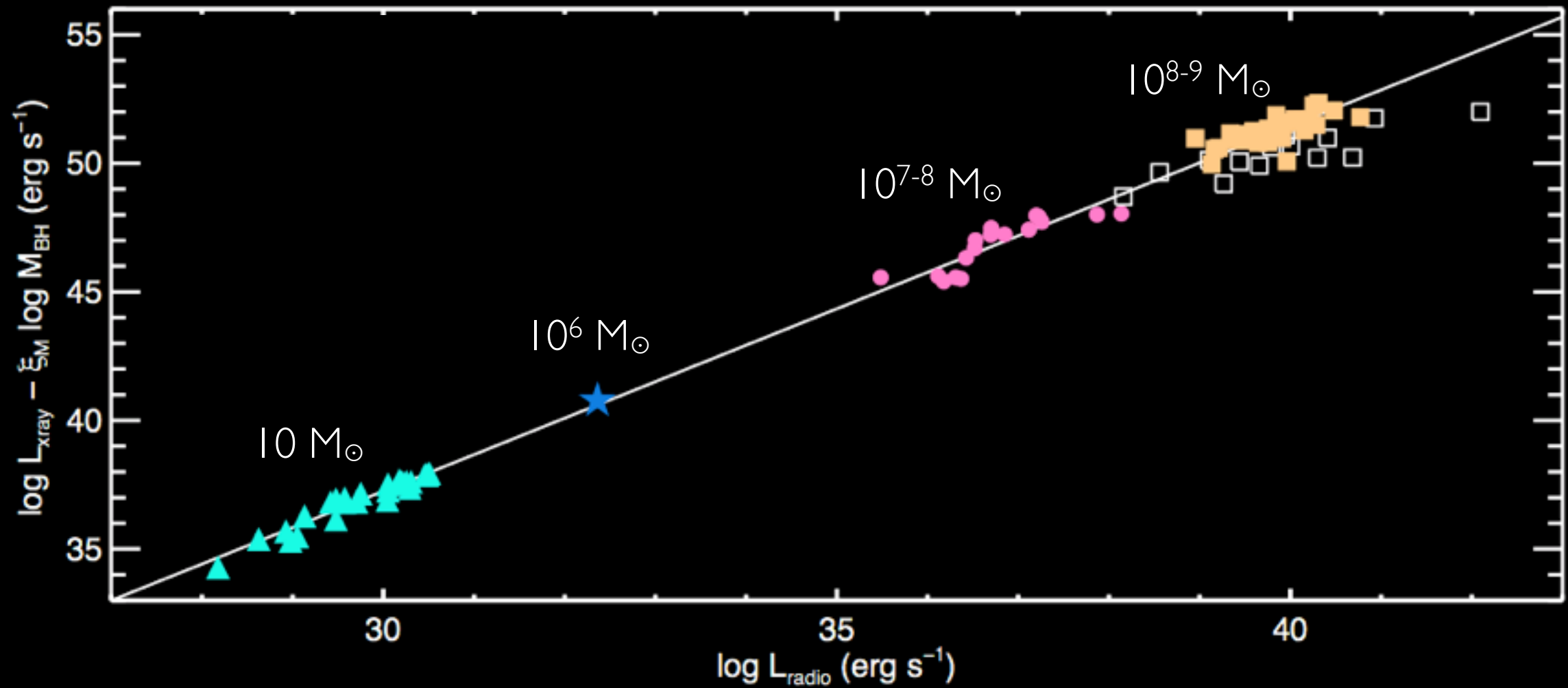
Wide Field Imager:

ΔE : 125 eV
Field of View: 40 arcmin
High countrate capability

Microquasars



Black hole L_X - L_R fundamental plane



Plotkin et al 2012